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# Light extinction by Secondary Organic Aerosol: an intercomparison of three broadband cavity spectrometers

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## Abstract

Broadband optical cavity spectrometers are maturing as a technology for trace gas detection, but only recently have they been used to retrieve the extinction coefficient of aerosols. Sensitive broadband extinction measurements allow explicit separation of gas and particle phase spectral contributions, as well as continuous spectral measurements of aerosol extinction in favourable cases. In this work, we report an inter-comparison study of the aerosol extinction coefficients measured by three such instruments: a broadband cavity ring-down spectrometer (BBCRDS), a cavity-enhanced differential optical absorption spectrometer (CE-DOAS), and an incoherent broadband cavity-enhanced absorption spectrometer (IBBCEAS). Experiments were carried out in the SAPHIR atmospheric simulation chamber as part of the NO<sub>3</sub>Comp campaign to compare the measurement capabilities of NO<sub>3</sub> and N<sub>2</sub>O<sub>5</sub> instrumentation. Aerosol extinction coefficients between 655 and 690 nm are reported for secondary organic aerosols (SOA) formed by the NO<sub>3</sub> oxidation of  $\beta$ -pinene under dry and humid conditions. Despite different measurement approaches and spectral analysis procedures, the three instruments retrieved aerosol extinction coefficients that were in close agreement. The refractive index of SOA formed from the  $\beta$ -pinene + NO<sub>3</sub> reaction was 1.61, and was not measurably affected by the chamber humidity or by aging of the aerosol over several hours. This refractive index is significantly larger than SOA refractive indices observed in other studies of OH and ozone-initiated terpene oxidations, and may be caused by the large proportion of organic nitrates in the particle phase. In an experiment involving ammonium sulphate particles the aerosol extinction coefficients as measured by IBBCEAS were found to be in reasonable agreement with those calculated using Mie theory. The results of the study demonstrate the potential of broadband cavity spectrometers for determining the optical properties of aerosols.

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## 1 Introduction

The contribution of suspended atmospheric particles to radiative forcing is both large and poorly constrained (Forster et al., 2007; Menon, 2004). The aerosol extinction – that is, the optical loss per unit length along a light path due to absorption and scattering of light by particles – is usually dominated by scattering. Nevertheless, absorption of light by particles is important in many contexts and has attracted much recent attention (Andreae and Gelencser, 2006; Bergstrom et al., 2007; Myhre, 2009; Martins et al., 2009; Jacobson, 1999; Alexander et al., 2008; Mellon et al., 2011). To better quantify the influence of atmospheric particles on radiative forcing, further characterization of their optical properties is necessary.

Current approaches for studying the optical properties of particles are mainly limited in two respects: Firstly, significant experimental artefacts and lack of agreement are associated with some aerosol absorption measurements (Kirchstetter et al., 2004; Andreae and Gelencser, 2006). Secondly, aerosol optical properties are often not available at ultraviolet wavelengths (where the absorption of carbonaceous aerosols sometimes increases strongly); this incomplete spectral information limits assessment of the aerosol influence on radiative balance and tropospheric photochemistry (Martins et al., 2009; Bergstrom et al., 2007; Hoffer et al., 2006; Andreae and Gelencser, 2006). Several studies have therefore called for new approaches to measuring the optical absorption of particulate matter in the atmosphere (Hallquist et al., 2009; Andreae and Gelencser, 2006; Kirchstetter et al., 2004).

Over the last decade, optical cavity methods have greatly advanced the characterization of aerosols (Pettersson et al., 2004; Moosmüller et al., 2009; Miles et al., 2011; Riziq et al., 2007). The high sensitivity of these methods results from the very long effective pathlengths – typically hundreds of metres to tens of kilometres – that are achieved inside high finesse optical cavities. Sensitive and accurate aerosol extinction coefficient measurements can also be combined with a separate measurement of the scattering coefficient in the so-called extinction-minus-scattering approach to re-

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trieve the aerosol absorption. Cavity ring-down spectroscopy (CRDS) has been used by several groups for laboratory, field, and airborne studies of aerosol optical properties (Smith and Atkinson, 2001; Moosmüller et al., 2005; Strawa et al., 2003; Thompson et al., 2002; Ma and Thompson, 2012). Extinction coefficient detection limits of well below  $1 \text{ Mm}^{-1}$  (equivalent to  $10^{-8} \text{ cm}^{-1}$ ) have been demonstrated with CRDS systems (Moosmüller et al., 2009). A related method, cavity attenuated phase shift (CAPS) spectroscopy, has also been used for quantifying aerosol extinction over a narrow wavelength band (Kebabian et al., 2007; Massoli et al., 2010). Although Miles et al. (2010) have recently demonstrated a CRDS system for measuring aerosol extinction spectra, typical CRDS and CAPS system are limited to one or two wavelengths; moreover, care must be taken to account for gas absorption when quantifying the aerosol extinction.

Richer spectral information can be acquired by combining optical cavities with broadband light sources (Ball and Jones, 2003; Bitter et al., 2005; Venables et al., 2006; Platt et al., 2009; Ruth et al., 2013). As with laser-based CRDS, this approach was first applied to trace gas detection but has since been extended to quantify aerosol extinction. Early work with broadband light sources and optical cavities focussed on single wavelength or passband measurements. Thompson and Spangler (2006) spectrally integrated the intensity of an incandescent tungsten bulb through an optical cavity to monitor aerosol extinction. Varma et al. (2009) used the fractional absorption of the  $\text{O}_2$  B-band to quantify the total extinction of the sample at 687 nm, and subsequently retrieved the aerosol extinction after removing the small contribution of gas phase absorption. Thalman and Volkamer (2010) similarly quantified the aerosol extinction at blue wavelengths based on the  $\text{O}_2$ -dimer absorption at 477 nm or the  $\text{H}_2\text{O}$  absorption at 443 nm. As demonstrated below, it is also possible to determine the aerosol extinction directly at an arbitrary wavelength if absorption by gases can be accounted for. Broadband optical cavity methods do not attain the sensitivity of typical CRDS and other laser-based cavity approaches, but crucially their spectral information allows absorbing gases to be quantified and their absorption contribution to the total sample

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extinction removed, thereby permitting the underlying aerosol extinction to be quantified. The method is also readily extended to different wavelength regions – particularly to shorter wavelengths where aerosol absorption often increases (Andreae and Gelencser, 2006; Kirchstetter et al., 2004; Chen and Bond, 2010; Hoffer et al., 2006; Martins et al., 2009). Very recently, several groups have started to exploit the potential of broadband cavity systems to achieve both spectrally continuous measurements of aerosol extinction and to perform measurements at short visible and ultraviolet wavelengths (Zhao et al., 2013; Washenfelder et al., 2013; Wilson et al., 2013).

In June 2007, three broadband optical cavity systems took part in an intercomparison study at the SAPHIR atmosphere simulation chamber. Although the primary focus of the campaign was to compare instrumental performance in measuring  $\text{NO}_3$  (Dorn et al., 2013) and  $\text{N}_2\text{O}_5$  (Fuchs et al., 2012), the optical extinction due to particles in the chamber was also retrieved by the broadband instruments. In this study, we report the aerosol extinction coefficients on four days of the campaign, and compare the instruments' performance and different analytical approaches used to retrieve the aerosol extinction coefficients. For the instrument with the most extensive dataset, we compare extinction coefficients measured for ammonium sulfate aerosols with Mie theory calculations and assess the instrument's long-term stability. Finally, we report for the first time the refractive index of the secondary organic aerosol (SOA) produced by  $\text{NO}_3$  oxidation of  $\beta$ -pinene.

## 2 Experimental

The aerosol extinction coefficient measurements reported here were made during the instrument intercomparison campaign "NO3Comp" held in June 2007 at the SAPHIR atmosphere simulation chamber, located in the Forschungszentrum Jülich, Germany (Bohn et al., 2005; Rohrer et al., 2005; Brauers et al., 2007; Apel et al., 2008). The campaign was designed for the intercomparison of instruments which measure  $\text{NO}_3$  (Dorn et al., 2013) and  $\text{N}_2\text{O}_5$  (Fuchs et al., 2012), and was subsequently broadened

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to include NO<sub>2</sub> measurements (Fuchs et al., 2010). Most of the instruments monitored NO<sub>3</sub> via its absorption band at 662 nm using either optical cavities or multipass DOAS. Some instruments also measured N<sub>2</sub>O<sub>5</sub> via its thermal dissociation to NO<sub>3</sub> in a second heated cavity channel. The three NO<sub>3</sub> cavity instruments that adopted a broadband approach were also sensitive to other gas phase absorbers and to aerosol extinction, as discussed in this work.

In this study, we report aerosol extinction coefficient measurements from 15–18, 20, and 21 June 2007 when aerosol was either injected into or produced inside the chamber. Oxidation pathways and SOA formation yields and composition from these experiments have been described previously (Fry et al., 2009, 2011; Rollins et al., 2009). The SAPHIR chamber was equipped with a comprehensive suite of instruments to monitor the gas phase composition inside the chamber, the chemical and physical properties of aerosols, and environmental variables such as temperature, pressure, and actinic flux. Particle size distributions were measured every 7 min using a scanning mobility particle sizer (SMPS, TSI 3936L85) and additionally with an aerosol mass spectrometer (Aerodyne TOF-AMS, 6 min acquisition time). SMPS size distributions were internally corrected for multiply charged particles. The chamber had been purged overnight with synthetic air prior to all experiments; the measurements therefore began in clean, particle-free air with particle number concentrations below 10 cm<sup>-3</sup>.

The three broadband cavity instruments deployed at the campaign were based on broadband cavity ring-down spectroscopy (BBCRDS, University of Leicester), cavity-enhanced differential optical absorption spectroscopy (CE-DOAS, University of Heidelberg), and incoherent broadband cavity-enhanced absorption spectroscopy (IB-BCEAS, University College Cork). The instruments shared a common measurement methodology: the spectrum of a broadband light source, transmitted through an optically stable cavity, was measured at a spectrograph either as a time-resolved ring-down signal (BBCRDS) or as a time-integrated intensity (CE-DOAS and IB-BCEAS). The instrument configurations and the spectral analysis procedures differed and are described in detail below and summarized in Table 1. To retrieve the aerosol extinction,

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absorption by NO<sub>3</sub>, H<sub>2</sub>O, NO<sub>2</sub> and O<sub>3</sub> was removed from the sample's total extinction by fitting and then subtracting the instruments' own measured spectral contributions from these gas-phase absorbers (Ball et al., 2004). The BBCRDS instrument was an extractive system located beneath the chamber, while the cavity of the IB-BCEAS instrument extended across the full 20 m length of the chamber in an open path configuration (see Fuchs et al., 2010; Varma et al., 2009). The CE-DOAS instrument was located on the floor inside the chamber and also had an open path cavity. Together, the instruments measured aerosol extinction at wavelengths between 655 and 690 nm (Table 1). The measurements reported in Sect. 3 are based on experiments when the aerosol extinction inside the chamber was above the instrumental detection limits. The experimental conditions for these days are given in Table 2. The intercomparison of aerosol extinction coefficients is restricted to the NO<sub>3</sub> +  $\beta$ -pinene experiments on 20 and 21 June when all three instruments were operating.

## 2.1 Broadband Cavity Ring-Down Spectroscopy (BBCRDS)

The BBCRDS technique is based on the simultaneous measurement of wavelength-resolved ring-down transients in a high finesse optical cavity containing the sample (Bitter et al., 2005; Ball and Jones, 2003; Leigh et al., 2010). Light from a pulsed broadband dye laser (662 nm emission maximum with FWHM of 16 nm) pumped at 20 Hz by a frequency-doubled Nd:YAG laser was directed into a 183 cm long ring-down cavity formed by two highly reflective mirrors (maximum reflectivity,  $R_{\text{max}} = 99.996\%$  at 680 nm). Although the BBCRDS instrument deployed at SAPHIR was optimized for detection of NO<sub>3</sub> via its 662 nm absorption band (Yokelson et al., 1994), absorption due to NO<sub>2</sub> and water vapour and aerosol extinction were also measured within the instrument bandwidth. To preserve the cleanliness of the mirrors' surfaces, the mirror mounts were purged by 0.5 L min<sup>-1</sup> of dry synthetic air, giving a length through the sample of 1.56 m. The measured extinction was corrected by a length factor,  $L_f$ , (cf. Eq. 1 below) to account for the sample being excluded from the mirror purge regions. Sample gas was drawn into the cavity (formed inside a 19 mm internal diameter Teflon tube) through

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four parallel Teflon tubes (40 cm long and 3 mm internal diameter) which protruded 15 cm inside the SAPHIR chamber to sample gas uncompromised by wall effects. The sample flow rate was  $10 \text{ L min}^{-1}$ , corresponding to a residence time of 2.7 s in the instrument. Light exiting the ring-down cavity was dispersed in wavelength by a spectrometer (0.36 nm FWHM spectral resolution) and imaged onto a clocked CCD camera (XCam CCDRem2). The time evolution of individual ring-down events was recorded simultaneously at 512 different wavelengths, corresponding to 512 clocked rows on the CCD camera. Typically, fifty ring-down events were integrated on the CCD camera before the image was read to a computer for processing and storage. Extinction spectra were integrated for 60 s.

The spectrum of the sample's extinction coefficient,  $\varepsilon(\lambda)$ , was calculated from sets of wavelength-resolved ring-down times measured when the cavity contained the sample,  $\tau(\lambda)$ , and when back-flushed with dry zero air,  $\tau_0(\lambda)$ :

$$\varepsilon(\lambda) = \frac{L_f}{c} \left( \frac{1}{\tau(\lambda)} - \frac{1}{\tau_0(\lambda)} \right) = \sum_i \alpha_i(\lambda) + \varepsilon_{\text{con}}(\lambda) \quad (1)$$

where  $L_f$  is the length factor,  $c$  is the speed of light,  $\alpha_i(\lambda) = \sigma_i(\lambda) \times N_i$  are the absorption coefficients of the various fitted molecular absorbers (with absorption cross section  $\sigma_i$  and number density  $N_i$ ), and  $\varepsilon_{\text{con}}(\lambda)$  is the wavelength-dependent extinction coefficient due to all unstructured contributions to the spectrum (mainly aerosol extinction). In accordance with Eq. (1), the extinction spectra were fitted by reference absorption spectra of  $\text{NO}_3$ ,  $\text{H}_2\text{O}$  and  $\text{NO}_2$  and a quadratic polynomial function to account for the unstructured contributions,  $\varepsilon_{\text{con}}(\lambda)$ . In addition to the aerosol extinction,  $\varepsilon_{\text{con}}(\lambda)$  includes (i) instrument artefacts (typically  $< 5 \text{ Mm}^{-1}$ ) due to (for example) degradation in the cavity's alignment caused by temperature variations, and (ii) any molecular absorptions that are unstructured or only very broadly structured over the BBRDS instrument's bandwidth. Ozone has a weak, but detectable, unstructured absorption, which was subtracted from  $\varepsilon_{\text{con}}(\lambda)$  using ozone concentrations measured by the SAPHIR chamber's core instruments; however we cannot exclude the possibility of small additional

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contributions from unknown molecular absorbers present in the complex gas mixtures generated within SAPHIR. The aerosol extinction coefficient measurements reported here from the BBRDS instrument are the mean values of  $\varepsilon_{\text{con}}(\lambda)$  computed over the wavelength range 655 to 665 nm, after subtraction of any ozone contribution.

## 2.2 Cavity-Enhanced Differential Optical Absorption Spectroscopy (CE-DOAS)

The CE-DOAS instrument comprised an optical cavity formed by two highly reflective mirrors ( $R = 99.9985\%$ ) separated by 0.62 m (Platt et al., 2009; Meinen et al., 2010). Light from a red LED, stabilised to  $27 \pm 2^\circ \text{C}$ , was focussed into the optical cavity. Light exiting the cavity through the second mirror was coupled into a 400 m quartz fibre (NA = 0.22) of 5 m length and guided out of the SAPHIR chamber. The fibre was attached alternately to a photomultiplier tube for cavity ring-down measurements or to a spectrograph for intensity measurements.

Ring-down measurements were used to determine the effective length of the light path in the cavity when the chamber (and thus the cavity) was free from aerosols and molecular absorbers in the detected wavelength range. A photomultiplier tube and high-speed current-to-voltage amplifier detected the ring-down signal, which was recorded by a high-speed digitizer. The mean optical path length  $L_0$  was calculated from the ring-down time. Cavity transmission spectra were recorded by coupling the fibre into an Ocean Optics USB2000 spectrometer with 1.06 nm spectral resolution. The spectrometer was temperature stabilized to  $0.0^\circ \text{C} \pm 0.1^\circ \text{C}$  and each spectrum was acquired over approximately 60 s.

During the campaign, the cavity of the instrument was mounted inside the SAPHIR chamber and placed 0.5 m above the ground to minimize wall effects. The cavity was located directly in the air flow of the chamber fan and the system was operated in an open path configuration. To avoid degradation of the mirror reflectivity during the measurement, a  $5 \text{ L min}^{-1}$  synthetic airflow was used to purge the mirrors. The single pass distance through the sample (that is, the distance between the mirrors excluding the purge length) was 0.50 m. With this configuration, an effective optical path length

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of about 8.5 km was attained when the chamber was free of aerosol and absorbing gases.

The spectral evaluation was based on the DOAS principle with additional corrections for the effective length of the light path in the cavity (Platt et al., 2009). A reference spectrum was recorded in the clean chamber (that is, aerosol and absorber-free) before the start of each day's experiment and applied to the trace gas and aerosol extinction retrieval. For trace gas analysis, the total measurement time was 5 to 6 min. Reference absorption cross sections of NO<sub>3</sub>, NO<sub>2</sub>, O<sub>3</sub> and H<sub>2</sub>O were fitted to the measured spectrum to retrieve the number densities of these species. To retrieve the aerosol extinction, small molecular absorptions from these species were subtracted from the total extinction over the wavelength interval from 673 to 677 nm. The acquisition time for extinction spectra was about 60 s. The derived trace gas column densities and absolute absorption were converted to concentrations and aerosol extinction coefficient using Eq. (2). According to Platt et al. (2009), the effective path length,  $L_{\text{eff}}$ , can be derived from the mean optical path length  $L_0$  derived from the ring-down measurement and applying a correction.

$$\varepsilon(\lambda) = \frac{\ln(I/I_0)}{L_{\text{eff}}} = \frac{\ln(I/I_0)}{L_0} \times \frac{(I/I_0 - 1)}{\ln(I/I_0)} = \frac{(I/I_0 - 1)}{L_0} \quad (2)$$

The statistical uncertainty ( $1\sigma$ ) for the aerosol extinction measurement was typically about 1 Mm<sup>-1</sup>. The aerosol extinction retrieval requires a stable intensity  $I_0$ ; however, the temperature stabilisation of the LED was insufficient during this set of experiments and gave rise to an estimated error of about 10 Mm<sup>-1</sup>.

### 2.3 Incoherent Broadband Cavity-Enhanced Absorption Spectroscopy (IBBCEAS)

The IBBCEAS instrument was customized for use in the SAPHIR chamber, with the optical cavity extending across the full length of the chamber (Varma et al., 2009; Fuchs

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et al., 2010). The cavity mirrors (reflectivity,  $R \sim 0.9984$  at 687 nm; -21 m radius of curvature; 40 mm diameter) were separated by 20.13 m, but purging the cavity mirrors to maintain their cleanliness reduced the sample length to 18.3 m  $\pm$  0.2 m. Light from a high radiance short-arc Xe arc lamp was coupled into the cavity, and light transmitted through the cavity was recorded with a monochromator/CCD system (Andor DV401). The instrument's spectral resolution was approximately 0.6 nm FWHM.

The following procedure was adopted on each day of the campaign. A reference intensity spectrum was measured for several minutes at the start of the day to acquire the baseline intensity spectrum,  $I_0$ , through the clean chamber. The reflectivity of the cavity mirrors was calibrated by inserting a loss-calibrated anti-reflection coated optical flat into the cavity. Subsequently, the cavity transmission spectrum was recorded every 5 s. The extinction coefficient of the sample,  $\varepsilon(\lambda)$ , is given by Fiedler et al. (2003):

$$\varepsilon(\lambda) = \sum_i \sigma_i(\lambda) N_i + \varepsilon_{\text{con}}(\lambda) = \left( \frac{I_0}{I} - 1 \right) \left( \frac{1 - R}{L_s} \right) \quad (3)$$

where the cavity's transmitted intensity spectra for the clean and sample-filled chamber are given by  $I_0(\lambda)$  and  $I(\lambda)$ , respectively.  $L_s$  is the sample length inside the cavity and  $R$  is the geometric mean reflectivity of the cavity mirrors. Number densities of H<sub>2</sub>O, NO<sub>3</sub>, and NO<sub>2</sub> were retrieved using a least squares fit of their respective absorption cross-sections to the measured absolute extinction. The small absorption of O<sub>3</sub> was corrected using the concentration measurements from the SAPHIR ozone monitor.

The aerosol extinction was retrieved based on the change in the fractional absorption of the O<sub>2</sub> B-band absorption at 687 nm. The fractional absorption changes because aerosol and other extinction processes shorten the effective optical path length through the sample (Varma et al., 2009; Thalman and Volkamer, 2010):

$$L_{\text{eff}} = \frac{L_s}{1 - R + \varepsilon L_s} \quad (4)$$

The B-band of O<sub>2</sub> is particularly appropriate for this analysis because the concentration and thus the absorption of O<sub>2</sub> are constant in the real atmosphere. Fortunately, absorp-

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tion by other atmospheric constituents is small at 687 nm, with maximum absorption of  $0.6 \text{ Mm}^{-1}$  (75 ppbv  $\text{NO}_2$ ) and  $0.6 \text{ Mm}^{-1}$  (180 ppbv  $\text{O}_3$ ) in these experiments. The underlying aerosol extinction coefficient,  $\varepsilon_A$ , can be retrieved directly from the differential absorption as described by Varma et al. (2009):

$$\varepsilon_A = \alpha_{\text{O}_2} \left( \frac{I}{\Delta I_{\text{O}_2}} \right) - \left( \frac{1-R}{L_s} \right) \quad (5)$$

where  $\Delta I_{\text{O}_2}$  is the difference between the baseline intensity and the intensity at the absorption maximum of the  $\text{O}_2$  band, and  $\alpha_{\text{O}_2}$  is the absorption coefficient of oxygen. We assume that the change in the total extinction (after subtracting the small contribution of gas phase absorption) arises from aerosol extinction, which is reported for 687.0 nm. Equation (5) notably does not require  $I_0$  to be known. As a result, the ratio ( $I/\Delta I_{\text{O}_2}$ ) remains stable even when the lamp intensity  $I$  fluctuates.

The rotational structure of the B-band is not resolved by the spectrometer. The question arises as to whether the unresolved fractional absorption of the band changes predictably with the aerosol extinction (Ball and Jones, 2003). To evaluate this issue, a synthetic, high resolution spectrum of the B-band (Jacquinet-Husson et al., 2011) was calculated for different aerosol extinctions in the cavity and then convolved with the IBBCEAS instrument function. These simulations confirmed that the fractional absorption of the unresolved band scaled linearly with the aerosol extinction within the range of aerosol extinctions reported in this study. Only when  $\varepsilon_A$  exceeded  $100 \text{ Mm}^{-1}$  did the fractional absorption deviate noticeably from linear behaviour. In a second test, the magnitude of the aerosol extinction retrieved using Eq. (5) agreed closely with that observed at 684 nm, a nearby wavelength that was largely free of molecular absorption and hence indicative of the aerosol extinction. This latter approach to estimating the aerosol extinction was highly susceptible to changes in lamp intensity, however, and is not considered further in this study.

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### 3 Results

We first compare the aerosol extinction coefficient of ammonium sulfate particles, measured by the IBBCEAS setup, against that calculated from Mie theory. The aerosol extinction coefficients retrieved by all three instruments are then compared in two experiments in which SOA was formed following the oxidation of  $\beta$ -pinene by  $\text{NO}_3$ . Figures also show concentrations of  $\text{NO}_3$  to indicate the chemical processes occurring during the experiments.

#### 3.1 Comparison against Mie theory

Varying concentrations of polydisperse ammonium sulfate particles were introduced into the chamber during the experiment on 15 June. While the main purpose of this experiment was to test the instruments' abilities to quantify  $\text{NO}_3$  and  $\text{N}_2\text{O}_5$  in the presence of ammonium sulphate aerosol, it also provided the opportunity to test the ability of the IBBCEAS instrument to quantitatively retrieve aerosol extinction for an aerosol sample of known concentration, particle size and composition. Ammonium sulfate,  $(\text{NH}_4)_2\text{SO}_4$ , has well characterised optical properties in the visible and is commonly used for calibrating light scattering instruments (Toon et al., 1976; Riziq et al., 2007; Moosmüller et al., 2009; Myhre et al., 2004). Figure 1 compares the extinction coefficient measured by the IBBCEAS instrument at 687 nm with that calculated from Mie theory assuming spherical particles (Wickramasinghe, 1973; Bohren and Huffman, 1983; Mätzler, 2002). The Mie calculations used the size distributions measured by the SMPS, and assumed that particles had a density of  $1.77 \text{ g cm}^{-3}$  and a refractive index of  $1.52 + 1 \times 10^{-7}i$  at about 680 nm (Toon et al., 1976). The real and imaginary parts of the refractive index vary little around these wavelengths. The maximum relative humidity (RH) in the experiment (ca. 60 % RH) remained below the deliquescence point of ammonium sulfate (80 % RH) and the influence of relative humidity was not considered in the Mie calculations (Martin, 2000).

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Reasonable agreement was found between the measured and calculated extinction coefficients. A slow upward drift of about  $0.7 \text{ Mm}^{-1}$  in the measured aerosol extinction is evident prior to introducing aerosol into the chamber. Aerosol was then introduced into the chamber in two steps between 10:45–11:25 LT and 12:30–14:30 LT. Two corresponding steps are evident in the measured extinction which broadly agree with the Mie theory calculation (Fig. 1), although the measured aerosol extinction is about 25 % lower at the extinction maximum ( $\sim 15:00$  LT) than that calculated from Mie theory. The difference of  $2 \text{ Mm}^{-1}$  at this point is not unreasonable, taking into account the precision of the measurements ( $0.2\text{--}0.4 \text{ Mm}^{-1}$ , see below), a calibration uncertainty of about  $0.6 \text{ Mm}^{-1}$ , and instrument drift of around  $1 \text{ Mm}^{-1}$ . In the latter case we note that the measured extinction coefficient is lower than the calculated coefficient by over  $1 \text{ Mm}^{-1}$  at 12:00 LT – that is, before the start of the second, main introduction of aerosol into the chamber. Any errors made in correcting for the absorption of gases would also influence the retrieved aerosol extinction values. Other factors that contribute to the observed difference are due to the uncertainty in the calculated extinction. The latter depends on uncertainties in the measured size distribution, the assumption of spherical particles, the value and extrapolation of the literature refractive index to 687 nm, and the dependence of the refractive index of ammonium sulfate aerosols on humidity (Riziq et al., 2007). There may also have been sampling uncertainties associated with the different locations of the SMPS and IBBCEAS instruments and potential spatial inhomogeneities in the chamber. Taking these several factors into account, we consider the agreement between the measurements and Mie calculations to be satisfactory and to broadly substantiate the reliability of the measured aerosol extinction values.

In the experiment of 16–17 June, SOA was formed from limonene oxidation under dry conditions (dew point  $< -44^\circ\text{C}$ ) and aged over a period of about 40 h (Fry et al., 2009). Figure 2 shows the extinction coefficient measured over the two day period. Limonene was introduced in two stages: in the first stage, 10 ppbv of limonene was added to the chamber followed by the addition of  $\text{NO}_2$  and  $\text{O}_3$ . Formation of  $\text{NO}_3$  and subsequent oxidation of limonene produced large numbers of small particles, which

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gradually grew to detectable sizes by around 12:00 LT. A second addition of  $\text{NO}_2$ ,  $\text{O}_3$ , and a further 10 ppbv of limonene around 14:30 LT increased particle sizes further, resulting in a much stronger increase in the aerosol extinction. The large increase in aerosol extinction following the second limonene addition likely arises through the condensation of oxidation products onto preexisting particles, which causes them to grow to sizes that scatter red light more efficiently. The maximum aerosol extinction coefficient of  $20 \text{ Mm}^{-1}$  was observed around 19:00 LT, some 4.5 h after the second introduction of limonene. The subsequent decrease in extinction broadly tracks the decreasing aerosol mass measured by the AMS and is attributable to dilution of the chamber contents and gradual deposition of particles to the chamber walls. Interestingly, the peak in the extinction occurs slightly later (and is rather broader) than the peak in the aerosol mass, presumably because the continued particle growth results in particles that more efficiently scatter red light and thus off-set the initial loss of aerosol mass. The chamber was flushed after 18:00 LT on 17 June, reducing particle concentrations to  $< 100 \text{ cm}^{-3}$  and leaving an effectively “clean” chamber by 24:00 LT. The SMPS was unfortunately not operating during this experiment and thus the refractive index of the SOA produced by  $\text{NO}_3$  + limonene oxidation chemistry could not be determined.

The 16–17 June experiment presented an opportunity to evaluate the long term stability of the IBBCEAS instrument, which operated over the entire duration of the experiment and relied on a single measurement of  $I_0$  (Eq. 2) on the morning of 16 June before the start of the experiment. The stability of the instrument can be gauged by a comparison of the extinction measured at the start and end of the experiment in the absence of aerosol, and which should be the same for a stable instrument. The extinction coefficient is marginally below zero (ca.  $2 \text{ Mm}^{-1}$ ) at the end of the measurements, indicating that the system was exceptionally stable over this extended period.

### 3.2 Intercomparison of aerosol extinction coefficient measurements

All three broadband cavity spectrometers measured the extinction coefficients of SOA formed following the  $\text{NO}_3$ -initiated oxidation of  $\beta$ -pinene under dry conditions (20 June)

and at 60 % RH (21 June). On both days the chamber was loaded with high  $\text{NO}_x$  and  $\text{O}_3$  levels to form  $\text{NO}_3$  at up to ca. 300 pptv (20 June) and ca. 120 pptv (21 June) (Rollins et al., 2009).  $\text{N}_2\text{O}_5$ , formed through the reaction of  $\text{NO}_3$  and  $\text{NO}_2$ , was also a large reservoir for  $\text{NO}_3$ . Subsequent addition of 15 ppbv  $\beta$ -pinene to the chamber resulted in a sharp drop in the  $\text{NO}_3$  concentration, which remained depleted for about an hour before recovering strongly. SOA mass concentrations of 30 to  $40\text{ }\mu\text{g m}^{-3}$  were formed during the experiments. No seed particles were added to the chamber on either day. The time series of the aerosol extinction coefficients from the three instruments are shown in Fig. 3a for 20 June and Fig. 4a for 21 June.

The precisions of the aerosol extinction coefficient measurements were evaluated before the formation of aerosol in the chamber over uninterrupted periods where each instrument exhibited stable and consistent performance (Table 3). On this basis, the precisions of the CE-DOAS, BBCRDS and IBBCEAS measurements were, respectively, 0.70, 0.37 and  $0.25\text{ Mm}^{-1}$  on 21 June for integration times of 60 s for CE-DOAS and BBCRDS and 5 s for IBBCEAS. Although the periods analysed were mostly longer than 1 h, the above values represent the short-term measurement-to-measurement variability. Significant instrument drift is also evident in the aerosol-free measurements on both days, however. To capture these effects, the mean and standard deviation of the instruments over the entire period before aerosol formation are also included in Table 3. Here, the mean extinction coefficients for the instruments fall between  $-1.2$  and  $2.4\text{ Mm}^{-1}$  (with the exception of the BBCRDS measurements on 20 June, which had a larger offset). The stability of this baseline is indicated by the standard deviation, which ranged from 2 to  $4\text{ Mm}^{-1}$  for the BBCRDS and CE-DOAS extinction coefficients measurements. These variations in the aerosol-free extinction measurements are less than 6 % of the observed extinction maxima, and demonstrate the importance of acquiring good reference measurements.

On 20 June (Fig. 3), the instruments exhibited very similar temporal behaviour with good agreement observed over the early stages of SOA formation and growth ( $\sim 09:10$ – $10:00$  LT). The extinction coefficients from the three instruments agree to within 10 %

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of their mean at the maximum. A change in the performance of the BBCRDS system is evident after approximately 13:30 LT, when occasional mistrigging of the camera with respect to the laser pulse caused the precision to deteriorate and thus the extinction measurements to become more scattered. Variations of up to  $7\text{ Mm}^{-1}$  in the responses of the BBCRDS and CE-DOAS instruments are evident in the particle-free chamber (that is, before introduction of  $\beta$ -pinene). In the case of the CE-DOAS measurements, these fluctuations arose from thermal instability of the LED. The same explanation probably applies to the observed undulations in the CE-DOAS aerosol extinction after 11:00 LT on 20 June.

The agreement between the extinction coefficients values is even better throughout the SOA experiment under humid conditions on 21 June (Fig. 4) and maximum extinction coefficients values agree to within 4 % of their mean. In the 4 h prior to particle formation, the extinction is very low ( $< 2.5\text{ Mm}^{-1}$  for BBCRDS and  $< \sim 0.5\text{ Mm}^{-1}$  for IBBCEAS). The non-zero CE-DOAS values (of up to  $9\text{ Mm}^{-1}$ ) over the same period are again attributed to temperature fluctuations of the LED.  $\text{NO}_2$  (80 ppbv) and  $\text{O}_3$  (90 ppbv) were added to the chamber at 09:24 and 09:36 LT. No obvious discontinuities in the measured aerosol extinction coefficients of any of the instruments coincide with  $\text{NO}_2$  and  $\text{O}_3$  injections into the chamber, indicating that the absorption of  $\text{NO}_2$  and  $\text{O}_3$  were properly accounted for in retrieving the aerosol extinction coefficients.

Correlation plots of the BBCRDS and CE-DOAS extinction coefficients against those from the IBBCEAS measurements are shown in Figs. 3b and 4b. The abscissa values are based on the IBBCEAS measurements due to its continuous dataset over both experiments and its demonstration of smaller means and instrument drift during the aerosol-free initial phases of the 20 and 21 June experiments. To account for the different sampling times, and to facilitate computation of the refractive index discussed below, the extinction coefficients are averaged over the 7 min sampling interval of the SMPS. As is clear from the time profiles for both days, the extinction coefficients from the three instruments are highly correlated. Table 4 summarises the parameters characterising the linear regressions of each correlation. For the data acquired on 20 June,

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the slopes of the correlation plots are similar for the BBCRDS (0.90) and CE-DOAS (0.88) systems, with respective correlation coefficients,  $R^2$ , of 0.983 and 0.978. Excellent agreement was found between the measurements on 21 June (Fig. 4b), with correlation coefficients of 0.998 (BBCRDS) and 0.991 (CE-DOAS). The slopes for the  
5 BBCRDS and CE-DOAS values are respectively 0.93 and 1.02, with small y-intercepts.

### $\beta$ -pinene SOA refractive index

The refractive index of particles formed by the  $\text{NO}_3$  oxidation of  $\beta$ -pinene was calculated from the measured extinction coefficients and aerosol size distributions for both experiments conducted under dry and humid conditions. The refractive index is  
10 a complex property ( $m = n + ik$ ) and both real and imaginary components are usually required. We have assumed here that the SOA generated in our experiments is non-absorbing (i.e.  $k = 0$ ) because the major constituents of monoterpene SOA – including carboxylic acids, aldehydes, oxy-aldehydes, oxy-carboxylic acids, dicarboxylic acids, and hydroxyl carboxylic acids – do not absorb at the red wavelengths of the  
15 present experiments (Kanakidou et al., 2005). Previous studies have found that similar monoterpene SOAs are non-absorbing (Kim et al., 2012; Nakayama et al., 2010; Kanakidou et al., 2005; Lang-Yona et al., 2010; Schnaiter et al., 2005). In the specific experiments discussed here, Fry et al. (2009) found that the SOA comprised a large fraction (up to 40–45 %) of organic nitrates, with possibly some carbonyl groups. Neither carbonyls nor organic nitrates absorb appreciably above 350 nm (Roberts and  
20 Fajer, 1989).

Refractive index values were calculated using measurements from each instrument over the half hour of the extinction maximum. The extinction measurements (equally weighted) of the individual instruments were averaged to give a single value for the  
25 extinction coefficient; the measured size distributions over this time period were also averaged. The refractive index computed for both experiments was 1.61. This value is notably larger than previous studies have reported for the refractive indices of monoterpene SOA (Table 5). Assuming a 10 % uncertainty in the mean extinction coefficient

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(which is representative of the agreement between the instruments and of their stated uncertainties), we estimate an associated uncertainty in the refractive index of  $\pm 0.03$  for both experiments. The time profile of the refractive index (not shown) showed no evolution in the value of the refractive index over the several hours of either experiment.

## 5 4 Discussion

Although the aerosol extinction measurements reported above share a common approach of measuring the spectrum of broadband light transmitted through an optical cavity, we stress that the measurement methodologies, calibration, and spectral analysis procedures of the systems differed appreciably. In particular, the measurement  
10 principle was either based on recording ring-down transients (BBCRDS), or on measuring the time-integrated intensity (CE-DOAS and IBBCEAS). Instrument calibration varied between spectrally-dependent ring-down times in zero air (BBCRDS and CE-DOAS) and the intensity attenuation by a calibrated low-loss optic (IBBCEAS). In the case of the IBBCEAS system, the aerosol extinction was quantified from the change in  
15 the fractional absorption of a molecular absorption band (Eq. 5), as opposed to direct quantification of the continuum extinction underlying the molecular absorption signals measured within the instrument bandwidth (Eq. 3).

For all instruments, the spectral analysis procedure quantified the aerosol extinction by subtracting retrieved (or otherwise known) molecular absorptions from the total measured extinction. The accuracy of the retrieved aerosol extinction obviously depends  
20 on the completeness and accuracy of quantifying and removing gas phase absorption. Sudden changes in the  $\text{NO}_2$ ,  $\text{O}_3$ ,  $\text{NO}_3$ , and  $\text{H}_2\text{O}$  concentrations during the experiments did not influence the retrieved aerosol extinctions, showing that the absorption of these species was properly accounted for by the three distinct analysis procedures (Dorn et al., 2013). Still, we note that the aerosol extinction is an upper limit as any unknown  
25 gas phase absorption would contribute to the (apparent) aerosol extinction. Because so few species absorb at the long visible wavelengths used in these experiments (and no

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unexplained structured features were present in any of the instruments' fitted spectra), such absorption is expected to be minimal in the extinction measurements reported here.

The extinction coefficients of  $\beta$ -pinene SOA measured by the three instruments are in good agreement in the two experiments on 20 and 21 June. The extinction measurements are at different wavelengths (687 nm, and mean over 655–660 nm and 673–677 nm): because the spectral dependence of scattering is expected to be weak over this 30 nm spread of wavelengths, we consider that these small wavelength differences do not affect the comparability of the measurements. Accurate aerosol extinction measurements are predicated on accurate quantification of the total sample extinction. As the respective  $\text{NO}_2$  and  $\text{NO}_3$  intercomparisons show, the concentrations of these species retrieved by the three broadband systems from the molecular absorptions agreed closely with several other techniques (Fuchs et al., 2010; Dorn et al., 2013). The total sample extinction measured by the three systems can therefore be viewed with confidence within the uncertainty limits of the measurements. Moreover, independent evidence for the reliability of the extinction measurements is found in the satisfactory agreement between the measured and calculated ammonium sulfate extinction (taking into account that the extinction was close to the instrument sensitivity).

The performance of the instruments is clearly suited to the study of SOA formation in simulation chambers as this work demonstrates for modest concentrations of organic precursor. Based solely on the precision of the extinction coefficient measurements, the respective limits of detection ( $3\sigma$ ) of the three systems were respectively 1.2, 2.1, and 1.2  $\text{Mm}^{-1}$  for the BBCRDS and CE-DOAS systems over 60 s, and for the IBBCEAS system over 5 s. Other recently developed broadband systems by Washenfelder et al. (2013) and Zhao et al. (2013) have reported somewhat lower precisions of around 0.2  $\text{Mm}^{-1}$  (corresponding to a detection limit of 0.6  $\text{Mm}^{-1}$ ). These figures represent the best case performance of the instruments and do not take into account any drift in the instruments' baselines. Such instrumental drift was generally a larger source of uncertainty and limited the accuracy at small aerosol extinctions. Engineering improve-

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ments since the  $\text{NO}_3\text{Comp}$  campaign, including more frequent re-calibration of the baseline spectrum, have produced performances much closer to the best case values (Kennedy et al., 2011). Zhao et al. (2013) report a long term stability of around 1  $\text{Mm}^{-1}$ , which is only slightly worse than the detection limit of their instrument. The detection limits we report are well sufficient for monitoring aerosol extinction in polluted atmospheres (for instance, mean aerosol extinction coefficients are 121  $\text{Mm}^{-1}$  in Atlanta, and over 300  $\text{Mm}^{-1}$  in Beijing) and are possibly low enough for measurements in pristine environments (Carrico et al., 2003; He et al., 2009). If necessary, the detection limits of these broadband systems could be improved by increasing the cavity length or by using higher reflectivity cavity mirrors. As a point of comparison, a recently-developed broadband aerosol extinction spectrometer using a multipass White cell had a higher detection limit of 33  $\text{Mm}^{-1}$  (albeit over a very wide spectral range of 250–700 nm) compared to the optical cavity instruments in this work (Chartier and Greenslade, 2012). It should be noted that detection limits of cavity ring-down systems are typically well below 1  $\text{Mm}^{-1}$  (Moosmüller et al., 2009); nevertheless, for the broadband instruments, the aerosol extinction is typically obtained in addition to the quantification of trace gases.

The question remains as to the wider application of broadband cavity methods to the optical properties of aerosols. In particular, are these methods sufficiently sensitive and accurate to retrieve the aerosol absorption using the extinction-minus-scattering approach? Many atmospheric aerosols have single scattering albedos (the ratio of scattering losses alone to the aggregate losses from both scattering and absorption) of 0.9 or greater, so extracting the aerosol absorption places stringent demands on the accuracy of both extinction and scattering measurements. A 10 % uncertainty in the aerosol extinction coefficients is representative for the instruments in this study, based on the good agreement of data and their stated uncertainties. Combined with the uncertainty in the measurements of scattering losses, the instruments would thus be suited to the “extinction-minus-scattering” approach for aerosols with appreciable absorption, although for studies interrogating aerosol properties in the laboratory one can use much higher particle concentrations than typically present in ambient air. Reduc-

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ing the uncertainty in the aerosol extinction would extend the utility of these methods to a wider range of aerosols. Most of the measurement uncertainty of the instruments arises from either the mirror reflectivity calibration or the baseline stability. A more accurate and frequent calibration procedure (as, for example, in Kennedy et al., 2011) could therefore extend their applicability to weakly absorbing aerosols.

A second consideration concerns the prospects for extending the approach to other spectral regions, and particularly to the ultraviolet where the optical properties of aerosols are poorly characterised. A major limitation of the fractional absorption approach used by the IBBCEAS system in this study is that this method is limited to known absorptions at a few wavelengths. Oxygen's absorption bands as used by the IBBCEAS instrument rapidly become weaker at wavelengths further into the visible, and so are probably unusable. However several other absorption bands of well-defined atmospheric gases exist and could be used instead. For instance, Thalmann and Volkamer (2010) have also used this approach based on the absorption of the O<sub>2</sub> dimer and H<sub>2</sub>O at 477 and 443 nm, respectively. More generally, broadband optical cavity methods have now been extended into the blue and near-UV region for trace gas detection (Hoch et al., 2012; Chen and Venables, 2011; Axson et al., 2011; Chen et al., 2011) and several groups have very recently demonstrated instruments for spectrally continuous measurements of aerosol extinction. Washenfeller et al. (2013) studied several scattering, weakly absorbing, and strongly absorbing aerosols between 360 and 420 nm and developed a procedure to retrieve the complex, wavelength-dependent refractive index of monodisperse aerosols over this spectral region. Wilson et al. (2013) reported the extinction coefficient from 320 to 410 nm of SOA formed by the photolysis of nitroaromatic compounds in an atmospheric simulation chamber. For field observations, Zhao et al. (2013) have developed an IBBCEAS system for the 445 to 480 nm spectral region and have demonstrated concurrent measurement of NO<sub>2</sub> concentration and aerosol extinction in a polluted urban environment. A key benefit of the Washenfeller, Wilson, and Zhao instruments, like the broadband spectrometers in the present work, is their ability to acquire wavelength-resolved spectra of the atmosphere extinction and

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hence the explicit separation of the structured gas phase and unstructured aerosol contributions. These recent studies underscore the value of sensitive, broadband extinction measurements to characterise the optical properties of atmospheric particles.

This work also reports the first measurement of the refractive index of monoterpene SOA produced primarily by NO<sub>3</sub> oxidation. The  $\beta$ -pinene mixing ratios (15 ppb) and the resulting aerosol mass concentrations (30 to 40  $\mu\text{g m}^{-3}$ ) in this work were low compared to other monoterpene SOA refractive index studies (Kim et al., 2010; Nakayama et al., 2010). Kim and Paulsen (2013) note that high aerosol mass loadings are likely to result in proportionally more high volatility species in the particle phase, and are thus less representative of ambient aerosols which have lower volatility constituents. The  $\beta$ -pinene SOA refractive index was  $1.61 \pm 0.03$  and did not depend on the humidity, nor did it evolve with aging of the aerosol over several hours. The absence of any humidity effect is consistent with the findings of Fry et al. (2009) that humidity did not influence the SOA yield, density, or proportion of alkyl nitrates in the particle phase for the 20 and 21 June experiments. Previous studies of the refractive index of SOA have not varied humidity outside of a narrow range and provide no point of comparison. Kim et al. (2010) also observed that the refractive index evolved little during ozonolysis of  $\alpha$ - and  $\beta$ -pinene, whereas it increased with aerosol mass concentration during photochemical oxidation of the same compounds. Reporting on the same experiments as us, Fry et al. (2009) found that the SOA mass yield decreased at longer times, suggesting volatilization from the particle phase. This effect was more marked in dry than humid conditions; nevertheless, we observed no associated change in the refractive index.

The  $\beta$ -pinene SOA refractive indices found in this study are strikingly large. Previous monoterpene SOA studies have reported long wavelength visible refractive indices between 1.38 and 1.56, with most values below 1.5 (Table 5; see also Kim and Paulsen, 2013). Our high refractive index values also fall outside the range of refractive indices of individual BVOC oxidation products reported by Redmond and Thompson (2011). Kim et al. (2012) found that the refractive indices of limonene and  $\alpha$ -pinene photooxidation SOA were higher at low HC / NO<sub>x</sub> ratios. The HC / NO<sub>x</sub> ratio in our experiments

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was about 2 ppb C ppb<sup>-1</sup> NO<sub>x</sub>, which is more than three-fold lower than the reaction conditions in Kim et al. (6.3–33 ppb C ppb<sup>-1</sup>). Nevertheless, the high refractive index values we report are in keeping with the trend in refractive index with HC / NO<sub>x</sub> ratios that they found. Kim et al. (2012) suggested that the high refractive index may be associated with either a higher proportion of organic nitrates or with oligomers in the aerosol. While we cannot exclude the possibility of oligomerisation in our experiments, the very high proportion (up to 45 %, Fry et al., 2009) of organic nitrates produced in these experiments suggests that organic nitrates are the likely cause for the high SOA refractive index.

A particle's refractive index strongly influences the nature of its interaction with light. High refractive index  $\beta$ -pinene SOA particles will scatter light more efficiently than prior SOA studies would suggest. For example, at a refractive index of 1.50 (which is typical of OH and ozonolysis SOA) the aerosol extinction of the particles in these experiments would be 35 % lower than we observed. Similarly, Kim and co-workers have calculated that changing the SOA refractive index from 1.4 to 1.5 alters radiative forcing by 12–19 % for non-absorbing aerosol (Kim et al., 2010; Kim and Paulsen, 2013). The results from this study and from Kim et al. (2012) suggest that the high refractive index seems particularly linked with the products of NO<sub>3</sub> oxidation chemistry or relatively high NO<sub>x</sub> levels. NO<sub>3</sub> chemistry occurs primarily at night (although it may also be appreciable in forest canopies during daylight), and recent work has found significant nocturnal particle formation driven by NO<sub>3</sub> oxidation of biogenic emissions (Rollins et al., 2012). The optical properties of such NO<sub>3</sub>-influenced organic aerosol, and their impact on radiative forcing, therefore warrant further study. For nocturnally-produced aerosol, in particular, the evolution of the particles' optical properties would need to be considered over several hours across the transition from night into the daylight conditions of the next day.

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**Table 1.** Operating parameters of the BBCRDS, CE-DOAS, and IBBCEAS instruments and the wavelength range of the reported aerosol extinction,  $\varepsilon_A$ .

Instrument	Cavity parameters*	Sampling time	Spectral range of $\varepsilon_A$	Uncertainty in $\varepsilon_A$
BBCRDS (extractive)	$L_p = 1.83$ m $L_s = 1.56$ m $L_f = 1.05$ $R = 99.996\%$ (680 nm)	60 s	655–665 nm	$0.4\text{--}4$ Mm $^{-1}$
CE-DOAS (open path)	$L_p = 0.62$ m $L_s = 0.50$ m $R = 99.9985\%$ (655 nm)	60 s	673–677 nm	$10$ Mm $^{-1}$
IBBCEAS (open path)	$L_p = 20.1$ m $L_s = 18.3$ m $R = 99.84\%$ (687 nm)	5 s	687.0 nm	10 % (relative)

\*  $L_p$  is the physical separation between the cavity mirrors,  $L_s$  is the length of the sample region in the cavity, and  $R$  is the mirror reflectivity at the specified wavelength. For the BBCRDS system, the length factor,  $L_f$ , is similar, but not identical, to the ratio  $L_p/L_s$ , probably owing to intrusion of sample gas into the purge region around the mirrors.

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**Table 2.** Description of NO3Comp experiments, and indicative values of the maximum aerosol extinction coefficient, for days on which aerosols were added to or formed in the chamber.

Date	Experiment description	Max $\varepsilon_a$
15 Jun	(NH <sub>4</sub> ) <sub>2</sub> SO <sub>4</sub> aerosol + NO <sub>3</sub>	$6$ Mm $^{-1}$
16,17 Jun	Limonene + NO <sub>3</sub> : SOA formation	$20$ Mm $^{-1}$
18 Jun	Isoprene + NO <sub>3</sub> : (NH <sub>4</sub> ) <sub>2</sub> SO <sub>4</sub> seed aerosol; SOA formation	$< 2$ Mm $^{-1}$
20 Jun	$\beta$ -pinene + NO <sub>3</sub> (dry conditions). SOA formation	$64$ Mm $^{-1}$
21 Jun	$\beta$ -pinene + NO <sub>3</sub> (60 % RH). SOA formation	$86$ Mm $^{-1}$

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**Table 3.** Analysis of the extinction coefficients measurements by the three instruments in aerosol-free conditions. Two datasets are analysed: (1) *stable operating conditions* refers to data when the baseline was stable for an extended time, and is indicative of the measurement precision of the instrument; (2) *all aerosol-free conditions* describes all the extinction data for each day and provides a measure of system accuracy and instrumental drift. Values are reported as the mean and standard deviation,  $1\sigma$ , of the aerosol extinction coefficient for  $N$  spectra at the native time resolutions of the instruments (60 s for BBCRDS and CE-DOAS, and 5 s for IBBCEAS).

Instrument	Period		$N$	Mean	$\sigma$
	Date	Time		( $\text{Mm}^{-1}$ )	( $\text{Mm}^{-1}$ )
<i>Stable operating conditions (aerosol-free)</i>					
BBCRDS	21/6:	06:40–07:26 LT	29	3.2	0.36
		08:20–10:33 LT	72	2.1	0.37
CE-DOAS	21/6:	06:13–07:55 LT	92	−1.5	0.70
IBBCEAS	20/6:	07:02–09:10 LT	1475	2.2	0.44
	21/6:	06:26–10:33 LT	2837	0.4	0.25
<i>All aerosol-free conditions</i>					
BBCRDS	20/6:	before 09:10 LT	48	7.5	2.2
	21/6:	before 10:33 LT	114	2.4	0.58
CE-DOAS	20/6:	before 09:10 LT	142	−1.2	3.2
	21/6:	before 10:33 LT	238	1.6	3.7
IBBCEAS	20/6:	before 09:10 LT	1475	2.2	0.44
	21/6:	before 10:33 LT	2837	0.4	0.25

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**Table 4.** Summary of the correlations of the aerosol extinction coefficients retrieved by BBCRDS and CE-DOAS against that from the IBBCEAS measurements. Data were averaged over the SMPS sampling time.

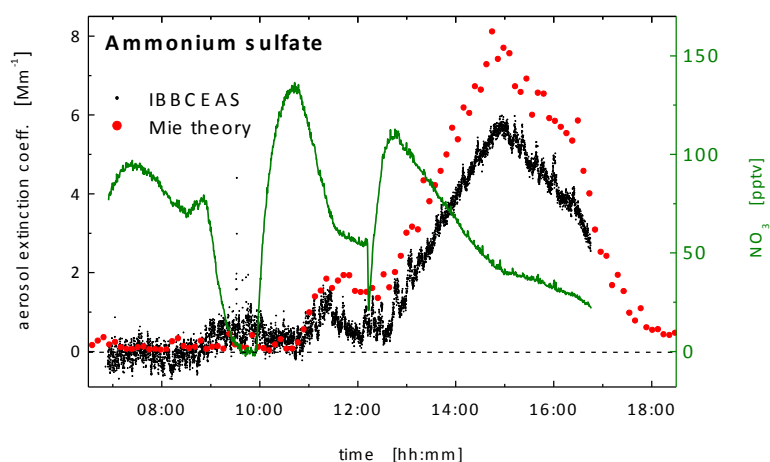
Instrument	Slope	y-intercept	$R^2$	$N$
20 Jun (dry)				
BBCRDS	0.90	$6 \text{ Mm}^{-1}$	0.9831	62
CE-DOAS	0.88	$-2 \text{ Mm}^{-1}$	0.9776	60
21 Jun (humid)				
BBCRDS	0.93	$2 \text{ Mm}^{-1}$	0.9983	47
CE-DOAS	1.02	$1 \text{ Mm}^{-1}$	0.9910	60

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**Table 5.** Refractive indices of SOA formed from monoterpenes.

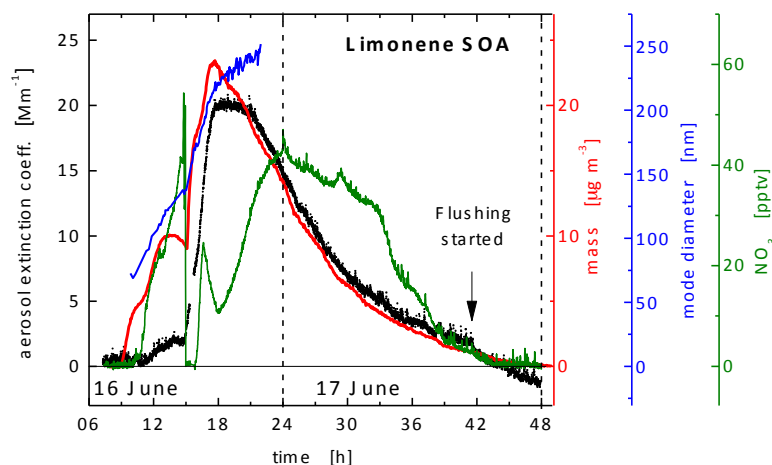
BVOC precursor	Reaction conditions	Refractive index, $\lambda$	Reference
$\alpha$ -pinene	Ozonolysis	1.5, 700 nm	Schnaiter (2005)
$\alpha$ -pinene	UV/O <sub>3</sub> /NO <sub>x</sub>	1.46, 700 nm	Yu (2008)
Holm oak emissions (98 % monoterpenes)	Photooxidation	1.53 ± 0.08, 532 nm	Lang-Yona (2009)
$\alpha$ -pinene	Ozonolysis	1.41, 532 nm	Nakayama (2010)
$\alpha$ -pinene	UV/NO <sub>x</sub>	} 1.4–1.5, 670 nm	} Kim (2010)
	Ozonolysis		
$\beta$ -pinene	UV/NO <sub>x</sub>	} 1.43–1.48, 670 nm	
	Ozonolysis		
$\alpha$ -pinene	Ozonolysis	1.49–1.51, 532 nm	Redmond (2011)
Limonene	UV/NO <sub>x</sub>	1.34–1.56, 532 nm	} Kim (2010)
$\alpha$ -pinene	UV/NO <sub>x</sub>	1.36–1.52, 532 nm	
$\beta$ -pinene	NO <sub>3</sub> oxidation	1.61 ± 0.03, 655–687 nm	This work

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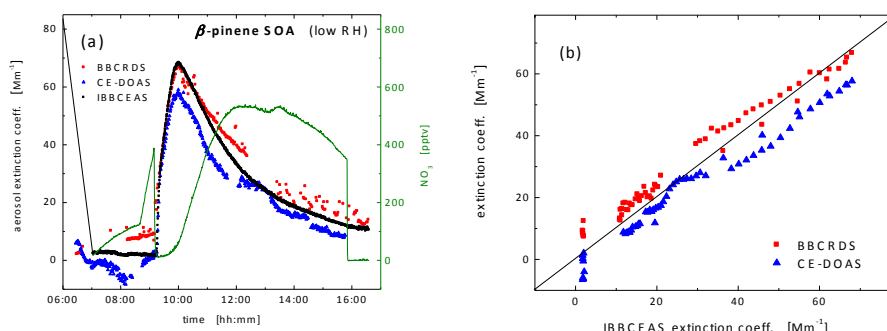
**Fig. 1.** Ammonium sulfate aerosol extinction coefficient on 15 June 2007 measured by the IBBCEAS instrument at 687 nm (black dots) and calculated from Mie theory (red dots) using the measured size distribution. Changes to the chamber composition occurred at: 08:55–09:52 LT (flushing of chamber with clean, humid air); 09:53 LT (addition of 80 ppbv O<sub>3</sub>); 09:56 LT (addition of 9 ppbv NO<sub>2</sub>); 10:45–11:25 LT (addition of 5  $\mu\text{gm}^{-3}$  of (NH<sub>4</sub>)<sub>2</sub>SO<sub>4</sub> aerosol); 12:12 LT (addition of 9 ppbv NO<sub>2</sub>); 12:30–14:30 LT (addition of 12  $\mu\text{gm}^{-3}$  of (NH<sub>4</sub>)<sub>2</sub>SO<sub>4</sub> aerosol).

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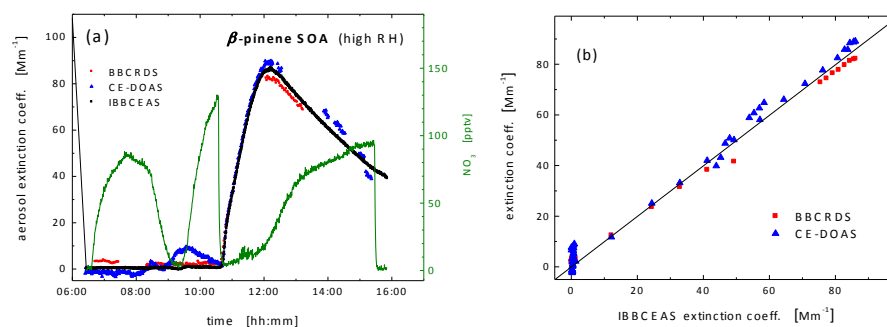
**Fig. 2.** IBBCEAS measurement of the extinction coefficient of SOA at 687 nm (black dots) formed by  $\text{NO}_3$  oxidation of limonene on 16–17 June 2007. Reactants were added to the chamber on 16 June at 06:20 LT (30 ppbv  $\text{NO}_2$ ), 06:26 LT (10 ppbv limonene), 10:30 LT (50 ppbv  $\text{O}_3$ ), and 14:30 LT (30 ppbv  $\text{O}_3$ , 30 ppbv  $\text{NO}_2$ , 10 ppbv limonene). Total aerosol mass (red) and mode diameter (blue, scaled) from the AMS are also shown. Continuous flushing of the chamber commenced shortly before 18:00 LT on 17 June, leaving a particle-free chamber by midnight.

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**Fig. 3.** (a) Time profile of the aerosol extinction coefficient following  $\text{NO}_3$  oxidation of  $\beta$ -pinene under dry conditions (20 June). Changes to the chamber composition occurred at: 06:27–06:37 (80 ppbv  $\text{NO}_2$ ), 06:50 LT (15 ppbv  $\text{O}_3$ ), 08:40 LT (80 ppbv  $\text{O}_3$ ), and 9:09 (15 ppbv  $\beta$ -pinene). (b) Correlation plots of the aerosol extinction coefficients retrieved by the BBCRDS and CE-DOAS instruments against those from the IBBCEAS measurements. The 1 : 1 line is shown as a solid line.

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**Fig. 4. (a)** Time profile of the aerosol extinction coefficients following  $\text{NO}_3$  oxidation of  $\beta$ -pinene under humid conditions (21 June). Changes to the chamber composition occurred at: 06:23–06:30 LT (160 ppbv  $\text{O}_3$ ), 06:35 LT (1 ppbv  $\text{NO}_2$ ), 08:26 LT (flushed chamber to 60 % RH), 09:24–09:34 LT (80 ppbv  $\text{NO}_2$ ), 09:36 LT (90 ppbv  $\text{O}_3$ ), and 10:34 LT (15 ppbv  $\beta$ -pinene). **(b)** Correlation plots of the aerosol extinction coefficients retrieved by the BBCRDS and CE-DOAS instruments against those from the IBBCEAS measurements. The 1:1 line is shown as a solid line.