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Authors	O'Leary, Deirdre M.;Ruth, Albert A.;Dixneuf, Sophie;Orphal, Johannes;Varma, Ravi M.
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# The near infrared cavity-enhanced absorption spectrum of methyl cyanide

Deirdre M. O'Leary<sup>a,1</sup>, Albert A. Ruth<sup>a,b,\*</sup>, Sophie Dixneuf<sup>a,b,2</sup>, Johannes Orphal<sup>c,3</sup>, and Ravi Varma<sup>a,4</sup>

<sup>a</sup> Physics Department, University College Cork, Cork, Ireland.

<sup>b</sup> Environmental Research Institute (ERI), University College Cork, Cork, Ireland.

<sup>c</sup> Laboratoire Interuniversitaire des Systèmes Atmosphériques (LISA), CNRS UMR 7683, 94010 Créteil Cedex, France.

\*Corresponding author: phone: +353-21-4902057, fax: +353-21-4276949, e-mail: a.ruth@ucc.ie

**Abstract** The absorption spectrum of methyl cyanide ( $\text{CH}_3\text{CN}$ ) has been measured in the near IR between 6000 and 8000  $\text{cm}^{-1}$  with a resolution of 0.12  $\text{cm}^{-1}$  using Fourier transform incoherent broadband cavity-enhanced absorption spectroscopy. The spectrum contains several only weakly perturbed spectral regions; potential vibrational combination bands contributing to the spectrum are outlined. Line positions and cross-sections of  $\text{CH}_3\text{CN}$  between 6814 and 7067  $\text{cm}^{-1}$  have been measured at high-resolution of 0.001  $\text{cm}^{-1}$  using diode laser based off-axis cavity-enhanced absorption spectroscopy. A total of 4630 new absorption lines of  $\text{CH}_3\text{CN}$  are identified in this region. A value for the self-broadening coefficient has determined to be  $(3.3 \pm 0.2) \times 10^{-3} \text{ cm}^{-1} \text{ mbar}^{-1}$  for one isolated line at 7034.171  $\text{cm}^{-1}$ . Several line series have been identified in these regions and an autocorrelation analysis performed with a view to aiding future assignments of the rotational-vibrational transitions.

**Keywords** Methyl cyanide, acetonitrile, Fourier transform cavity-enhanced spectroscopy, broadband absorption, high resolution.

## 1. Introduction

Methyl cyanide ( $\text{CH}_3\text{CN}$ , see Fig. 1) is a simple organic nitrile (acetonitrile), which, as a polar aprotic solvent is of common interest in industrial processes; e.g. for the determination of pesticide residues [3]. In the atmosphere methyl cyanide occurs as a trace constituent mainly as a result of biomass burning [4]–[6] and, to an extent, due to industrial processes and traffic [6]–[8]. It has been detected by remote sensing [9]–[13] and *in-situ* in surface air [14], in the troposphere [10], [15], and in the stratosphere [16]–[18] with typical mixing ratios ranging from the low pptv region in the upper stratosphere [13], [19], [20] to many ppbv in the vicinity of bushfires [15]. It is generally accepted that

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<sup>1</sup> Present address: St. Luke's Radiation Oncology Network, St. Luke's Hospital, Rathgar, Dublin 6, Ireland.

<sup>2</sup> Present address: Forschungszentrum Jülich GmbH, IEK-8 (Troposphere), 52425 Jülich, Germany.

<sup>3</sup> Present address: Karlsruhe Institute of Technology, Institute for Meteorology and Climate Research (IMK), 76344 Eggenstein-Leopoldshafen, Germany.

<sup>4</sup> Present address: Physics Department, National Institute of Technology Calicut, Calicut, 673601 Kerala, India.

CH<sub>3</sub>CN has a long lifetime in the atmosphere (> 6 months [10]). Removal of CH<sub>3</sub>CN is mainly based on reaction with OH, but also oceanic losses and washout by rain have been suggested as effective sinks of atmospheric CH<sub>3</sub>CN [10], [12], [17].

Methyl cyanide is also of considerable importance in astrophysics. The main isotopic species was first detected in the interstellar medium by Solomon *et al.* in 1971 [21]. The main isotopologue and other isotopic species have since been observed in numerous regions in the interstellar medium [22]–[26], in comets [27], [28], and in nearby galaxies [29], as well as in the atmosphere of Saturn’s largest moon, Titan [30]–[34] (e.g. substantial mixing ratios of  $(1.4 - 1.5) \times 10^{-6}$  for CH<sub>3</sub>CN were reported on Titan below 1050 km by the Ion Neutral Mass Spectrometer aboard the Cassini mission). In diverse astrophysical topics [35]–[38] methyl cyanide is deemed a good spectroscopic probe of physical conditions such as temperature and density. One of the reasons for this significant interest in methyl cyanide by astronomers is the fact that it was one of the first molecules investigated by microwave spectroscopy and it is therefore one of the best studied symmetric top molecules. It has been the subject of a large number of laboratory studies to measure line positions and to determine spectroscopic assignments in the infrared. It has been also used as a probe molecule for special spectroscopic techniques such as Beam Maser spectroscopy [39], [40] and Laser Stark spectroscopy [41]:

Methyl cyanide is a prolate symmetric top molecule of C<sub>3v</sub> symmetry (see Fig. 1). It has 12 normal modes of vibration; four non-degenerate vibrations of *a*<sub>1</sub>-symmetry and four doubly-degenerate vibrations of *e*-symmetry. All vibrational modes are infrared and Raman active. Its ground state rotational structure has been studied in the far infrared between 7 and 60 cm<sup>-1</sup> [42]. 1400 pure rotational transitions were identified, and assignments up to *J* = 90 and *J* = 79 were obtained in the ground and *v*<sub>8</sub> = 1 states respectively. Pavone *et al.* [43] investigated the ground state of CH<sub>3</sub>CN by CO<sub>2</sub> laser difference frequency spectroscopy in the region 0.5 – 1.5 THz. Fundamentals, overtones and combination bands at low energies have been measured at high resolution in the mid infrared [44]–[49]. The 2*v*<sub>8</sub> overtone band was studied using Fourier transform spectroscopy (FTS) [44] where 1000 lines could be assigned. Antilla *et al.* [45] studied the *v*<sub>5</sub> fundamental, the *v*<sub>7</sub> + *v*<sub>8</sub> combination band and the *v*<sub>7</sub> + *v*<sub>8</sub> - *v*<sub>8</sub> hot band to obtain the ground state constants. Tolonen *et al.* [46] assigned 4000 lines to the *v*<sub>4</sub>, *v*<sub>7</sub> and 3*v*<sub>8</sub> bands. Paso *et al.* [47] took FTS data in the region between 1240 and 1650 cm<sup>-1</sup> and assigned 3900 lines to the *v*<sub>6</sub> and *v*<sub>7</sub> + *v*<sub>8</sub> bands. High resolution (0.0025 cm<sup>-1</sup>) measurements of the *v*<sub>4</sub> and *v*<sub>7</sub> fundamental bands of CD<sub>3</sub>CN are described in Ref. [48]. The *v*<sub>1</sub> and *v*<sub>5</sub> fundamental bands of CH<sub>3</sub>CN were measured by Huet [49] between 2370 and 3500 cm<sup>-1</sup>. 1700 lines were assigned to these bands. The high relevance of CH<sub>3</sub>CN is also demonstrated by two recent publications on FT absorption spectroscopy by Harrison and Bernath [50] and Allen *et al.* [51], who measured air-broadened infrared absorption spectra of CH<sub>3</sub>CN in the region 880 – 1700 cm<sup>-1</sup> and around 3 μm with a resolution of 0.015 cm<sup>-1</sup>. In [50], [51] cross-sections (for a range of temperatures; 207 – 297 K) were established with a multi-pass cell with a maximum path-length of 19.3 m.

The study of methyl cyanide has also recently been extended into the near infrared region by Rinsland *et al.* [52]. This study reports absorption cross-sections and integrated band intensities for spectra of CH<sub>3</sub>CN measured in the laboratory. Spectra were recorded at a resolution of 0.112 cm<sup>-1</sup>, at three temperatures in the wavenumber range 600 –

6500  $\text{cm}^{-1}$ . The  $\nu_4$  parallel band ( $\sim 920 \text{ cm}^{-1}$ ) was also comprehensively studied by Rinsland *et al.* [53]. In Ref. [53] the room temperature line positions and intensities for over 700 individual lines are reported together with nitrogen- and self-broadening coefficients as well as pressure induced line shift coefficients. In 2008 methyl cyanide has been added for the first time to the high-resolution transmission molecular absorption database (HITRAN, a compilation of molecular spectroscopic parameters for the simulation of transmission and emission of light in the atmosphere, [www.cfa.harvard.edu/hitran](http://www.cfa.harvard.edu/hitran)) [54]. Based on data in Ref. [53] and supplemented by unpublished data [53] a total of 3572 absorption features are now available in HITRAN. Parameters on self- and air-broadened line widths have also been provided in Ref. [53].

High resolution absorption measurements at even higher ground state energies ( $>6500 \text{ cm}^{-1}$ ) are experimentally challenging due to increasingly weak transitions in this spectral region. In this work high resolution absorption data of  $\text{CH}_3\text{CN}$  in the ground state have been measured beyond 6000  $\text{cm}^{-1}$  using two different cavity-enhanced approaches:

- (a) *Fourier transform incoherent broadband cavity-enhanced absorption spectroscopy* (FT-IBBCEAS) using a supercontinuum light source [55], [56] to establish overview spectra at medium resolution between 6000 and 8000  $\text{cm}^{-1}$ .
- (b) *Off-axis cavity-enhanced absorption spectroscopy* (CEAS) [57] using an external cavity diode laser to establish a list of line positions and intensities in the near infrared region of the spectrum between 6814 and 7067  $\text{cm}^{-1}$ .

The new absorption data of  $\text{CH}_3\text{CN}$  presented here are useful for the development of theoretical models of the ground state potential of  $\text{CH}_3\text{CN}$  and also potentially useful for future observation of  $\text{CH}_3\text{CN}$  using near infrared spectroscopy in astronomical studies.

## 2. Experimental

### 2.1. Off-axis cavity enhanced absorption setup

The experimental off-axis CEAS set-up is essentially the same as the one described in [58], [59]. The length of the optical cavity (Layertec mirrors:  $r = -2 \text{ m}$ ,  $R \sim 0.997$ ) was 113 cm and the volume enclosed by the sample chamber was approximately 2  $\text{dm}^3$ . An off-axis alignment of the optical cavity was employed to reduce the effect of the cavity's mode structure on the transmitted intensity. The incident beam entered the cavity at a distance of approximately 8 mm from the optical axis. The laser was typically scanned mode-hop free over 2  $\text{cm}^{-1}$  per spectrum. The signal transmitted by the cavity was detected by an auto-balanced InGaAs photoreceiver (detector Nirvana 2017, New Focus) and averaged over 2000 sweeps per spectrum. This resulted in a  $1\sigma$  noise level of  $1.02 \times 10^{-25} \text{ cm}^2 \text{ molecule}^{-1}$  and a maximum signal-to-noise ratio of approximately 375. More than 200 spectra were concatenated to cover the overall spectral range reported.

### 2.2. Fourier transform incoherent broadband cavity-enhanced absorption setup

The experimental FT-IBBCEAS set-up is essentially the same as the one described in [56]. The optical cavity was attached to a vacuum chamber consisting of a long stainless steel pipe (100 mm diameter) which was pumped by a molecular turbo pump to approximately  $1 \times 10^{-5}$  mbar before introducing the sample gas. The length of the optical cavity (Layertec mirrors:  $r = -50$  m,  $R \sim 0.999$ ) was 644 cm. A supercontinuum (SC) fibre laser (Fianium, SC450, 6W) was used as light source. This pulsed source operates at 80 MHz delivering pulses of  $\sim 5$  ps duration. Despite the higher brightness and better coupling efficiency of the SC source, due to low divergence and good spatial coherence in comparison to an arc lamp or light emitting diode, it was found that there was virtually no improvement of the signal to noise ratio when compared to the ‘super quiet’ Xe arc lamp (Hamamatsu) used in [56]. Even though the noise of the SC source was not quantified, the pulse-to-pulse intensity fluctuations of the source as a function of wavelength must be substantial. This fact is particularly severe because the light exiting the cavity was coupled into a FT spectrometer (Bruker, Vertex 80) supplied with a InGaAs photodiode sensor. In combination with an FT detection scheme the noise of the light source is fully transferred to the detector as opposed to an approach where light dispersion is used to determine the wavelength. Data were recorded at a resolution of  $0.12 \text{ cm}^{-1}$ .

### 2.3. Wavelength calibration and accuracy

The presence of a large number of  $\text{H}_2\text{O}$  absorption features due to minute quantities of water vapour in the sample chamber allowed for the absolute calibration of the wavenumber scale over a wide spectral region. The line positions of transitions in water [60] were taken from the HITRAN database [54]. The uncertainty of the positions is quoted as being between  $0.001$  and  $0.01 \text{ cm}^{-1}$ . In regions containing an insufficient number of water absorption lines, separate measurements were performed with a small quantity of a calibrating gas,  $\text{CO}_2$  or  $\text{NH}_3$ , in the CEA chamber. The positions of the absorption lines of  $\text{CO}_2$  [61] were also obtained from the HITRAN database [54]. The uncertainty in the positions of  $\text{CO}_2$  lines is quoted as being between  $0.0001$  and  $0.001 \text{ cm}^{-1}$ . The positions of  $\text{NH}_3$  lines, taken from [58], had an uncertainty of  $0.001 \text{ cm}^{-1}$ . Absorption by  $\text{CH}_3\text{CN}$  and the calibration gas was recorded simultaneously thereby enabling the absolute wavenumber calibration of  $\text{CH}_3\text{CN}$  lines over the entire spectral region covered. Calibration was performed by position interpolation over the whole wavenumber scale and subsequent linear regression. The standard deviation of the fit parameters enabled an estimation of the uncertainty of the line positions at between  $0.005 \text{ cm}^{-1}$  and  $0.01 \text{ cm}^{-1}$ .

### 2.4. Calibration of the mirror reflectivity

In order to determine absorption coefficients and cross-sections with CEAS, it is necessary to know the reflectivity of the cavity mirrors over the whole range of measurement. The reflectivity of the mirrors is determined by measuring the absorption of a known amount of  $\text{CO}_2$ , whose line intensities are well known. The HITRAN database lists line intensities,  $i_{\text{H}}$ , [ $1/(\text{cm} \times \text{molecule} \times \text{cm}^{-2})$ ] as the integral of the absorption cross-sections at a temperature of 296 K.

The reflectivity,  $R$ , may be obtained from the line intensity  $i_H$  as follows:

$$R = 1 - \frac{N d i_H}{\int \left( \frac{I_0(\tilde{\nu})}{I(\tilde{\nu})} - 1 \right) d\tilde{\nu}} \quad (1)$$

where  $N$  [molecule  $\times$  cm<sup>-3</sup>] is the number density of the calibration gas (CO<sub>2</sub>),  $d$  [cm] is the cavity length, and  $I$  and  $I_0$  are the transmitted intensities with and without the calibration gas, respectively. The absorption by CO<sub>2</sub> was examined between 6870 and 6990 cm<sup>-1</sup> for the CEAS experiment and between ~6460 and ~7760 cm<sup>-1</sup> for the FT-IBBCEAS experiment. The error associated with other parameters such as the length of the cavity, the pressure of the sample and the integrated area under an absorption peak (denominator in (1)) is small so that the uncertainty in  $R$  is dominated by the error in the line intensities of the CO<sub>2</sub> lines. This error is quoted by HITRAN to be between 5% and 10%. The associated error in the value of  $(1-R)$  was therefore estimated to be 10%. A graph of the mirror reflectivity spectra obtained from the CO<sub>2</sub> measurements for the two sets of mirrors used in this work is presented in Fig. 2.

## 2.5. Materials and Gas preparation

Pressures of 4-5 mbar of CH<sub>3</sub>CN was chosen to enable the observation of weak absorption features while still trying to minimize the effect of line broadening, and associated blending of absorption features, in the observed spectrum. The acetonitrile sample (Labscan Analytical Sciences, Spectroscan grade) had a quoted purity of 99.9% (GC assay). The maximum impurity level of water in the solvent was stated at 0.05%. Degasification of the CH<sub>3</sub>CN sample was carried out by performing a number of ‘freeze-pump-thaw’ cycles.

## 2.6. Water absorption feature and their elimination

The presence of water inside the cavity and in the beam path outside the cavity resulted in the presence of unwanted water absorption features in the spectra which potentially completely obscure any CH<sub>3</sub>CN absorption. Rigorous water removal strategies were therefore necessary. Prior to experiments, the stainless steel chambers in both experiments were cleaned extensively with HPLC grade acetone. The sample chambers were evacuated for several days using a turbo molecular pump before the commencement of the absorption measurements. Although no significant leakage of air into the evacuated sample volume was noted, the ‘growth’ of water absorption features was observed with time, which was caused by desorption of water vapour from the walls of the sample chamber. In order to remove the water in the lab air outside the cavity it was necessary to place a plastic cover around the experiments, enclosing the optical path where possible. The volume inside of the plastic cover was purged with either dried compressed air or nitrogen gas. In case of the FT-IBBCEAS experiments the FT spectrometer was constantly flushed with dried and filtered lab air. In the case of the CEAS experiments purge gases were passed through a P<sub>2</sub>O<sub>5</sub> trap to further remove any traces of water. Despite these measures water features could only be dramatically reduced but not entirely suppressed experimentally. The presence of the large number of water absorption lines (see section III.A.)

was thus used for direct calibration of the wavenumber scale. It should be noted, that only water absorption lines that are unaffected by saturation and arising from water inside the cavity were used to calibrate the wavenumber scale.

### 3. Results and discussion

#### 3.1. FT-IBBCEAS overview spectra

Fig. 3 (black trace) shows an overview spectrum of  $\text{CH}_3\text{CN}$  (4 mbar) between 6000 and 8000  $\text{cm}^{-1}$ , measured with a resolution of 0.12  $\text{cm}^{-1}$  in an acquisition time of 60 min. The data file can be found in the electronic supplementary material. The absorption lines in this region are extremely weak and have not been addressed experimentally before in the gas phase. Several regions of strongly overlapping vibrational combination bands occur in the spectrum, whose structure is not rotationally resolved. In some cases distinct P- or R-branches can be identified. The strongest band in the entire spectrum occurs between 6000 and 6150  $\text{cm}^{-1}$ , where the first harmonic overtones  $2\nu_1$  and  $2\nu_5$  can occur as well as  $\nu_1+\nu_5$ . This region represents a "classic" perpendicular band of a symmetric top molecule where the strongest features correspond to individual Q-branches (also see [52], [53]). A list of potentially occurring vibrational states in the region between  $\sim 6150$  and 7500  $\text{cm}^{-1}$  is given in Table I for the combination of three fundamental modes only (Note: the values in Table I are purely combinatorial and do not take the anharmonicity of the molecule into account – they must not be read as assignments in this region). Based on the absorption strength of the fundamentals [62], [63], the combinations that would be expected to contribute the strongest to the observed spectrum are highlighted in Table I. Due to the strong coupling of vibrational modes and the high congestion of the spectrum, a comprehensive analysis of the data cannot be easily performed. It is interesting, however, that the spectrum behaves relatively "normal" in this region and does not exhibit an overly complex (e.g. polyad) structure. An approach to the analysis based on anharmonic force field calculations [64] rather than harmonically coupled local modes, as presented by Ahmed and Henry [65] for measurement in the liquid phase in this region, may already describe the spectrum obtained at a satisfactory level. Above 7500  $\text{cm}^{-1}$  a significant decrease in the  $\text{CH}_3\text{CN}$  absorption is observed (see Fig. 4). Between 7500 and 8000  $\text{cm}^{-1}$  only the absorption of states consisting of a combination of at least four vibrational quanta contribute to the spectrum. The spectrum is highly congested and exhibits several complex vibrational band systems that cannot be resolved with the present setup. Overtone combination bands above 8000  $\text{cm}^{-1}$  have only been addressed at low resolution in the liquid phase, where inhomogeneous broadening dominates the line profiles [65], [66].

The spectrum in Fig. 3 also contains absorption lines that are due to  $\text{H}_2\text{O}$ . The corresponding lines can be identified on basis of the HITRAN database [54] (blue trace in Fig. 3 shown for comparison). The spectrum also contains (rotationally resolved) absorption features based on the impurity HCN. The HCN feature is based on the CH-stretch overtone  $2\nu_3$  with band centre at  $\approx 6520$   $\text{cm}^{-1}$  [67]. The HCN band in the spectrum was the only feature whose relative absorption appeared to vary time dependently from measurement to measurement. Its occurrence

could never be entirely avoided. The green trace in Fig. 3 shows the region which was additionally recorded with high-resolution off-axis CEAS (see Fig. 5) – it will be discussed in section III.B.

Fig. 3 and 4 demonstrate the advantage of combining IBBCEAS with a Fourier transform detection scheme. A large spectral region can be covered at a resolution which is relevant for air-broadened lines under normal conditions. The  $1\sigma$  noise based on data at  $7625\pm 25\text{ cm}^{-1}$  is approximately  $1\times 10^{-8}\text{ cm}^{-1}$  (60 min acquisition time). Despite the significant spectral and temporal noise of the supercontinuum source the signal-to-noise (SNR) ratio above  $7500\text{ cm}^{-1}$  is still  $\sim 60$ ; in the region around  $6000\text{ cm}^{-1}$ , where the absorption signal is close to saturation, the SNR can approach  $10^5$ .

### 3.2. Off-axis CEAS high resolution spectra

The cavity-enhanced absorption spectrum of methyl cyanide, recorded at a pressure of 5 mbar is shown in Fig. 5. Approximately 200 individual, overlapping spectral segments were concatenated to cover a spectral range of approximately  $250\text{ cm}^{-1}$ . The recorded spectrum of  $\text{CH}_3\text{CN}$  is very dense – 4630 new absorption transitions of methyl cyanide were identified corresponding to an average density of 18 lines per  $\text{cm}^{-1}$ . The rotational structure of the spectrum is in many parts not fully resolved. The largest absorption signal was detected at  $6930.962\text{ cm}^{-1}$ . The absorption cross-section of this line is  $6.54 \times 10^{-23}\text{ cm}^2$  with a corresponding absorption coefficient of  $8.14 \times 10^{-6}\text{ cm}^{-1}$ . The lowest detectable peak was measured at  $6914.257\text{ cm}^{-1}$  with an absorption cross-section of  $5.18 \times 10^{-25}\text{ cm}^2$  and an absorption coefficient of  $6.45 \times 10^{-8}\text{ cm}^{-1}$ . A comprehensive list of line positions and absorption strengths of  $\text{CH}_3\text{CN}$  obtained using CEAS is given in the supplementary material.

Since the spectrum of methyl cyanide was measured at 5 mbar, the effect of pressure broadening (see section III.C.) causes adjacent lines to blend into one another. Therefore zero-absorption was not always well defined. In each of the measured spectral segments a program was used to locate the positions of absorption minima which were assumed to be zero. A polynomial was then fitted through all of these positions to determine the zero-absorption baseline over the entire spectral segment.

Despite the efforts to remove water from the spectrum (see section II.F.) it was still necessary to identify the strongest  $\text{H}_2\text{O}$  lines by comparison with the HITRAN 2004 database and remove them from the  $\text{CH}_3\text{CN}$  absorption spectrum. As an example the spectral interval between  $6990$  and  $7000\text{ cm}^{-1}$ , measured with off-axis CEAS and FT-IBBCEAS, is shown in Fig. 6. The water features (represented in form of a blue "stick spectrum" based on the HITRAN integrated line data in this region) have been removed from the spectrum. Fig. 6 illustrates the importance of minimizing and removing the water absorption features from the spectrum.

#### *Spectral analysis*

The possible overtone and combination bands for absorption transitions, up to three quanta, in the region between  $6800$  and  $7100\text{ cm}^{-1}$  are listed in Table I. Since intensities of overtones and combination bands strongly decrease with increasing number of quanta [68], the bands considered in Table I only consist of three quanta. All vibrational



transitions in Table I are symmetry allowed. The overview spectrum in Fig. 5 was (somewhat arbitrarily) subdivided into three regions which will be discussed in turn:

*Region (i): 6814 – 6900 cm<sup>-1</sup>*

There are approximately 20 absorption transitions per cm<sup>-1</sup> in this region and several spectral series provide evidence of band structure. Fig. 7(a) shows an example of a series of 13 absorption lines between 6875 and 6885 cm<sup>-1</sup> with equidistant energy separation of 0.66 cm<sup>-1</sup>. As a prolate symmetric top molecule (moment of inertia  $I_c = I_b > I_a$ , cf. insert Fig. 5) CH<sub>3</sub>CN may display equidistant line series for modes of motion where centrifugal distortion is negligible. The line separation  $\Delta\tilde{\nu} = 2B$  depends on the rotational constant,  $B$ , which was stated as 0.3068 cm<sup>-1</sup> in [49]. The value for  $2B$  matches the separation of the strong lines in Fig. 7(a) by  $\approx 0.05$  cm<sup>-1</sup>. An autocorrelation of the corresponding spectral region between  $\sim 6876$  and 6885 cm<sup>-1</sup> was performed and is exemplified in Fig. 7(b). The autocorrelation  $A_c(\zeta)$  of the wavenumber-dependent absorption coefficient,  $\alpha(\tilde{\nu})$ , is a mathematical tool to identify repeating patterns in the spectrum.  $A_c$  is essentially the cross-correlation of  $\alpha(\tilde{\nu})$  with itself at a lag  $\zeta$  [69]:

$$A_c(\zeta) = \int_{-\infty}^{\infty} \alpha(\tilde{\nu})\alpha(\tilde{\nu} - \zeta) d\tilde{\nu}. \quad (2)$$

The wavenumber differences,  $\Delta\tilde{\nu}$ , for the two most significant peaks obtained from the autocorrelation correspond to 0.68 cm<sup>-1</sup> and 1.30 cm<sup>-1</sup> (vertical arrows in Fig. 7(b)). These two values (No. 1 and 2 in Table II) can be considered fundamental and second harmonic spacings of the series, respectively. The error in the positions of the peaks was estimated empirically by fitting a number of Gaussian profiles to the background-subtracted peaks of the autocorrelation curve. The FWHM of the Gaussian closest to the maximum of the autocorrelation peak was used as an estimate of the error. Other reasonably undisturbed line series were identified in the spectrum and a list of dominant equidistant separations based on similar analyses in other spectral regions is shown in Table II. In Table II the values for  $\Delta\tilde{\nu}$  obtained from the autocorrelation are compared with those determined from repeating patterns of lines in the spectrum. Additionally the position of each line in Fig. 7(a) was plotted versus a corresponding (relative) quantum number  $J$ , which was arbitrarily assigned to each position. A linear regression was applied to the data as shown in Fig. 7(b) (red line) in order to retrieve an estimate of  $2B$  based on the assumption of negligible higher order vibration-rotation interactions. The quadratic nature of the residuals of the linear fit exhibit the effect of centrifugal distortion on the rotational energy levels as outlined in Fig. 7(c). The maximum error in the peak positions is  $\sim 0.01$  cm<sup>-1</sup> and lies within the symbol size of data points in Fig. 7(c). The observed deviations in the residuals are therefore significant and real. A second order polynomial was used as a fit function corroborating the fact that this series is weakly distorted by the presence of vibration-rotation interactions. The linear regression approach was used in the following analysis in other parts of the spectrum in order to retrieve values for  $\Delta\tilde{\nu}$  and indeed other wavenumber intervals of interest. For instance, a similar series of lines was identified in the region between 6865 and 6871 cm<sup>-1</sup> (Table II).

*Region (ii): 6900 – 7005 cm<sup>-1</sup>*

This region of the spectrum contains the lines of greatest absorption cross-section. The density of absorption features in this region is approximately 7 transitions per cm<sup>-1</sup> based on a peak counting approach. Potential vibrations occurring in this region of the spectrum are the combination bands  $\nu_4 + 2\nu_5$ ,  $2\nu_1 + \nu_7$  and  $\nu_1 + \nu_5 + \nu_7$ . These transitions have *e*-symmetry and result in perpendicular bands. The selection rule  $\Delta K = \pm 1$  for perpendicular bands allows *Q* branches to be spectrally displaced relative to one another. The separation of successive *Q* branches in different perpendicular bands of a given symmetric top molecule may vary considerably [68]. The band shape is strongly dependent on the vibration-rotation (Coriolis) interaction, hence the band appearance can vary from regular *P*, *Q*, *R* branches structure to structureless. In the region considered here no regular *P*, *Q*, *R*-band structures or obvious band centres were observed. Features consisting of many closely spaced absorption transitions amalgamated together are prevalent and probably correspond to band heads of *Q* branches of perpendicular bands. An example of such a feature is given in Fig. 8. Six of these strongly blended absorption features were noted in the region between 6975 and 7005 cm<sup>-1</sup>. They are approximately equidistantly spaced with a mean separation of  $5.4 \pm 0.1$  cm<sup>-1</sup>, suggesting that these features form part of a spectral series which is governed by the rotational *A* constants, which is known to be 5.27 cm<sup>-1</sup> [53].

*Region (iii): 7005 – 7067 cm<sup>-1</sup>*

This region of the spectrum is quite densely populated with an average of ca. 18 absorption transitions per cm<sup>-1</sup>. Considering up to three quanta, the only vibration that can occur in this region is the  $2\nu_5 + \nu_7$  combination band. The region exhibits distinct spectral patterns particularly in the range between 7020 and 7060 cm<sup>-1</sup>. Seemingly unperturbed spectral patterns are readily identifiable in this region. For example, a pattern of 32 “double lines”, separated by  $\approx 0.2$  cm<sup>-1</sup> is repeating at intervals of approximately 0.63 cm<sup>-1</sup> in the region between 7020 and 7040 cm<sup>-1</sup>. Values corresponding to an autocorrelation of the spectrum between 7027 and 7032 cm<sup>-1</sup> is shown in Table II. The four respective entries (No. 5...8) correspond to the four possible line separations of two adjacent double lines. The two values of the repetition interval (denoted a and b in the third column of Table II) were calculated by considering both lines of the doublet. A second series of seemingly unperturbed double lines is identified in the region between 7048 and 7060 cm<sup>-1</sup>. The two strongly absorbing transitions are displaced by approximately 0.28 cm<sup>-1</sup> relative to one another. An autocorrelation of the spectral region between 7049 and 7054 cm<sup>-1</sup> was obtained showing three significantly broadened features (Table II, No. 9...11). The values  $\Delta\tilde{\nu}$  correspond to the separation of the double lines, the repetition interval of the pattern, and the sum of both. In this pattern the second line of the doublet lay approximately in the centre of the repetition interval. For this reason, two estimates of the doublet separation and the repetition interval (labeled c, d, e and f respectively in Table II) were obtained by examining the spectrum directly. This is also the reason why the peaks in the autocorrelation plot appear significantly broadened.

Unfortunately, spectral information in the region between 7040 and 7049 cm<sup>-1</sup> was obscured owing to the presence of water vapour. It is conceivable that the origin of the vibrational combination band,  $2\nu_5 + \nu_7$ , lies within this region. Indeed the observed spacing of the repeating patterns at lower energies is greater than that observed for the patterns

at higher energies, suggesting that the observed lines belong to the corresponding *P* and *R* branches respectively.

Finally, in the regions between 7031 and 7034  $\text{cm}^{-1}$  and 7058 and 7060  $\text{cm}^{-1}$  series of lines were found in the spectrum consisting of line pairs where the separation between the lines gradually diminishes with increasing wavenumber.

### 3.3. Pressure broadening at 7034.17 $\text{cm}^{-1}$

The high density of absorption transitions in this spectrum makes it difficult to locate isolated lines for accurate measurements of line widths and pressure broadening parameters. An absorption transition at 7034.171  $\text{cm}^{-1}$  was chosen for a line width and pressure broadening analysis, because it is reasonably well isolated from adjacent lines. Self-pressure broadening of methyl cyanide was investigated by fitting Voigt profiles to this line measured at different pressures of  $\text{CH}_3\text{CN}$  between 0.5 and 10 mbar. An example at 5 mbar is presented in the insert of Fig. 9 together with an associated Voigt profile fitted to the experimental data. The Voigt fit provided values of 0.01263  $\text{cm}^{-1}$  and 0.0144  $\text{cm}^{-1}$  for the Gaussian ( $v_G$ ) and Lorentzian FWHM ( $v_L$ ) values, respectively. There was no evidence of collisional narrowing – the absorption profiles are reasonably well reproduced by Voigt profiles – although such an effect may be masked by noisy data in this case. The theoretically expected Doppler FWHM was calculated to be 0.0135  $\text{cm}^{-1}$ . The Lorentzian line width of the absorption line at 7034.171  $\text{cm}^{-1}$  as a function of  $\text{CH}_3\text{CN}$  pressure is shown in Fig. 9. A linear fit to the data yielded a self-broadening coefficient of  $3.28 \times 10^{-3} \text{ cm}^{-1} \text{ mbar}^{-1}$  with an intercept of  $-1.44 \times 10^{-3} \text{ cm}^{-1}$  (dashed line in Fig. 9). The main source of error leading to the small but non-zero intercept results from the contribution of lines next to 7034.171  $\text{cm}^{-1}$  which affects the fit, particularly for larger pressures. In order to account for line blending effects from adjacent absorption transitions, Voigt profiles were also fitted to the adjacent lines during the fitting procedure. However, the uncertainty due to line overlap is significantly larger in comparison to the error in the pressure reading ( $\pm 0.01 \text{ mbar}$ ) and the inherent fit uncertainty, which is typically less than 5%. A second linear fit constrained to go through the origin was applied to the data (red straight line in Fig. 9) and yielded a self-broadening coefficient of  $3.07 \times 10^{-3} \text{ cm}^{-1} \text{ mbar}^{-1}$ . The difference between the two values of the self-broadening coefficient obtained was taken as the uncertainty in the value of the broadening coefficient. The self-broadening coefficient was therefore measured to be  $(3.3 \pm 0.2) \times 10^{-3} \text{ cm}^{-1} \text{ mbar}^{-1}$ . This value for the self-broadening coefficient compares well to results obtained by Rinsland *et al.* who measured self-broadened half width coefficients as large as 2.0  $\text{cm}^{-1} \text{ atm}^{-1}$  [53].

## CONCLUSIONS

Fourier transform incoherent broadband cavity-enhanced absorption spectroscopy (FT-IBBCESA) has been applied to measure the absorption spectrum of methyl cyanide ( $\text{CH}_3\text{CN}$ ) between 6000 and 8000  $\text{cm}^{-1}$  with a resolution of 0.12  $\text{cm}^{-1}$ . The high energy region of the  $\text{CH}_3\text{CN}$  spectrum ( $>6500 \text{ cm}^{-1}$ ) has been addressed spectroscopically for the first time. Additionally, diode laser based off-axis cavity-enhanced absorption spectroscopy has been used to measure

line positions and cross-sections between 6814 and 7067  $\text{cm}^{-1}$  with high resolution of 0.001  $\text{cm}^{-1}$ . A total of 4630 new absorption lines of  $\text{CH}_3\text{CN}$  were identified in this region. The spectrum contains several only weakly perturbed spectral regions and potential vibrational combination bands contributing to the spectrum have been identified. Several line series have been identified in these regions and an autocorrelation analysis performed with a view to aiding future assignments of the rotational-vibrational transitions. The self-broadening of one isolated line at 7034.171  $\text{cm}^{-1}$  has been studied and the corresponding broadening coefficient was determined as  $(3.3 \pm 0.2) \times 10^{-3} \text{ cm}^{-1} \text{ mbar}^{-1}$ .

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#### APPENDIX

The data files of the overview spectra shown in Figs. 3 and 5 are available in the supplementary material together with a list of resolved absorption lines in the region measured by off-axis CEAS.

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# TABLES

**Table 1** Possible overtone and combination bands (three quanta) of CH<sub>3</sub>CN in the region between 5900 and 7500 cm<sup>-1</sup>. Values have been calculated using the frequencies of the fundamental vibrations of CH<sub>3</sub>CN [62] and are only indicative of potential contributions in the spectrum.

Combination Band	Symmetry	Wavenumber [cm <sup>-1</sup> ]
2v <sub>1</sub>	a <sub>1</sub>	~5908
v <sub>1</sub> + v <sub>5</sub>	e	~5963
2v <sub>5</sub>	a <sub>1</sub>	~6018
v <sub>2</sub> + v <sub>4</sub> + v <sub>5</sub>	e	~6196
v <sub>1</sub> + v <sub>2</sub> + v <sub>7</sub>	e	~6262
2v <sub>1</sub> + v <sub>8</sub>	e	~6270
v <sub>2</sub> + v <sub>5</sub> + v <sub>7</sub>	a <sub>1</sub>	~6317
v <sub>1</sub> + v <sub>5</sub> + v <sub>8</sub>	a <sub>1</sub>	~6325
2v <sub>5</sub> + v <sub>8</sub>	e	~6379
v <sub>2</sub> + v <sub>3</sub> + v <sub>5</sub>	e	~6661
v <sub>1</sub> + v <sub>2</sub> + v <sub>6</sub>	e	~6669
v <sub>2</sub> + v <sub>5</sub> + v <sub>6</sub>	a <sub>1</sub>	~6724
3v <sub>2</sub>	a <sub>1</sub>	~6801
2v <sub>1</sub> + v <sub>4</sub>	a <sub>1</sub>	~6828
v <sub>1</sub> + v <sub>4</sub> + v <sub>5</sub>	e	~6883
v <sub>4</sub> + 2v <sub>5</sub>	e	~6938
2v <sub>1</sub> + v <sub>7</sub>	e	~6949
v <sub>1</sub> + v <sub>5</sub> + v <sub>7</sub>	e	~7004
2v <sub>5</sub> + v <sub>7</sub>	e	~7059
2v <sub>1</sub> + v <sub>3</sub>	a <sub>1</sub>	~7293
v <sub>1</sub> + v <sub>3</sub> + v <sub>5</sub>	e	~7348
2v <sub>1</sub> + v <sub>6</sub>	e	~7356
v <sub>3</sub> + 2v <sub>5</sub>	a <sub>1</sub>	~7403
v <sub>1</sub> + v <sub>5</sub> + v <sub>6</sub>	a <sub>1</sub>	~7411
2v <sub>5</sub> + v <sub>6</sub>	e	~7446
v <sub>1</sub> + 2v <sub>2</sub>	a <sub>1</sub>	~7488
2v <sub>2</sub> + v <sub>5</sub>	e	~7543

**Table 2** Relevant recurring wavenumber differences in spectral regions between 6876 and 7054 cm<sup>-1</sup> identified by autocorrelation (Column 2) of the absorption signal in the corresponding region and by visual inspection of the spectrum (Column 3).

No.	Spectral region	Autocorrelation $\Delta\tilde{\nu}$ [cm <sup>-1</sup> ]	Observed series $\Delta\tilde{\nu}'$ [cm <sup>-1</sup> ]
1	6876 – 6885	0.68 ± 0.04	0.66 ± 0.02
2	6876 – 6885	1.30 ± 0.06	1.32 ± 0.04
3	6865 – 6871	0.51 ± 0.04	0.57 ± 0.02
4	6865 – 6871	1.22 ± 0.04	1.13 ± 0.04
5	7027 – 7032	0.18 ± 0.06	0.195 ± 0.005
6	7027 – 7032	0.44 ± 0.06	0.430 ± 0.007
7	7027 – 7032	0.62 ± 0.07	0.626 ± 0.005 <sup>a</sup> 0.627 ± 0.006 <sup>b</sup>
8	7027 – 7032	0.80 ± 0.05	–
9	7049 – 7054	0.26 ± 0.05	0.26 ± 0.03 <sup>c</sup> 0.33 ± 0.02 <sup>d</sup>
10	7049 – 7054	0.58 ± 0.06	0.600 ± 0.003 <sup>e</sup> 0.610 ± 0.003 <sup>f</sup>
11	7049 – 7054	0.95 ± 0.04	–

## FIGURE CAPTIONS

**Fig. 1.** Structure and molecular coordinate system (a, b, c) of methyl cyanide. This symmetric top molecule has moments of inertia  $I_c = I_b (= 9.122845 \times 10^{-46} \text{ m}^2 \text{ kg [1]}) > I_a (= 5.308074 \times 10^{-47} \text{ m}^2 \text{ kg [2]})$ .

**Fig. 2.** Reflectivity spectrum of mirrors used for FT-IBBCEAS (upper traces) and off-axis CEAS (lower traces). Trace without error bars (blue): Mirror reflectivity measured with double beam absorption spectrometer (Lambda1050, Perkin-Elmer). Solid and hollow dots: reflectivity values established via calibration with CO<sub>2</sub> at 13.3 mbar and 4 mbar, respectively. Trace with error bars (red): Data obtained with the double beam spectrometer scaled to the average of the CO<sub>2</sub> calibration data (these traces were used as the mirror reflectivity). Error bars correspond to the 1 $\sigma$  standard deviation of the respective calibration data (the point in brackets was left out).

**Fig. 3.** Left axis (Black trace): Cavity-enhanced absorption spectrum of CH<sub>3</sub>CN between 6000 and 8000 cm<sup>-1</sup> measured by FT-IBBCEAS with a resolution of 0.12 cm<sup>-1</sup> at a pressure of 4 mbar (data file in supplementary electronic material). Acquisition time 60 min. (Green trace): Region addressed by high resolution off-axis CEAS – section III.B. Right axis (Blue trace): "Stick spectrum" of integrated cross-sections of H<sub>2</sub>O according to the HITRAN database [54].

**Fig. 4.** Absorption spectrum of CH<sub>3</sub>CN at 4 mbar in the region between 7500 and 8000 cm<sup>-1</sup> measured at 0.12 cm<sup>-1</sup> resolution using FT-IBBCEAS. Only absorption to states consisting of four vibrational quanta contributes to the spectrum in this region. SNR is approximately 60.

**Fig. 5.** Off-axis cavity-enhanced absorption spectrum of CH<sub>3</sub>CN between 6814 and 7067 cm<sup>-1</sup> measured at a pressure of 5 mbar. There are ~4630 new absorption transitions in this spectrum; water features were removed. Approximately 200 individual overlapping spectral segments were concatenated to cover the spectral range shown (data file in supplementary electronic material).

**Fig. 6.** A 10 cm<sup>-1</sup> section of the CH<sub>3</sub>CN spectrum illustrating the importance of the removal of water absorption features. Upper trace (left axis): Fourier transform detection, resolution 0.12 cm<sup>-1</sup>. Lower trace (left axis): off-axis CEAS detection, resolution 0.001 cm<sup>-1</sup>. Stick spectrum (right axis, blue): H<sub>2</sub>O reference values obtained from the HITRAN database [54].

**Fig. 7.** Upper panel (a): Arbitrarily  $J$ -numbered series of seemingly equidistant absorption lines in the region between 6865 and 6871 cm<sup>-1</sup>. Middle panel (b): Autocorrelation analysis of the spectral region between 6876 and 6885 cm<sup>-1</sup>.

The two peaks highlighted by vertical arrows can be related to wavenumber differences in the spectrum. Lower panel (c): Peak positions versus relative (arbitrary) values of  $J$  with associated linear (red) and 2<sup>nd</sup> order polynomial (blue) fits. Lower panel: residuals of linear (red) and 2nd order polynomial fits (blue).

**Fig. 8.** One of several unresolved band heads in the spectral region around  $6993 \text{ cm}^{-1}$ .

**Fig. 9.** Lorentzian FWHM of the absorption line at  $7034.17 \text{ cm}^{-1}$  as a function of pressure. Dashed line: Linear regression. Red line: linear regression forced through origin. Insert: Example of a fit of Voigt function to the absorption line measured at 5 mbar resulting in  $\text{FWHM}_{\text{Gauss}}=0.0126 \text{ cm}^{-1}$  and  $\text{FWHM}_{\text{Lorentz}}=0.0144 \text{ cm}^{-1}$  and residuals in the lower panel of the insert.

### Supplementary material

Data shown in Fig. 3

Data shown in Fig. 5

Figure 1

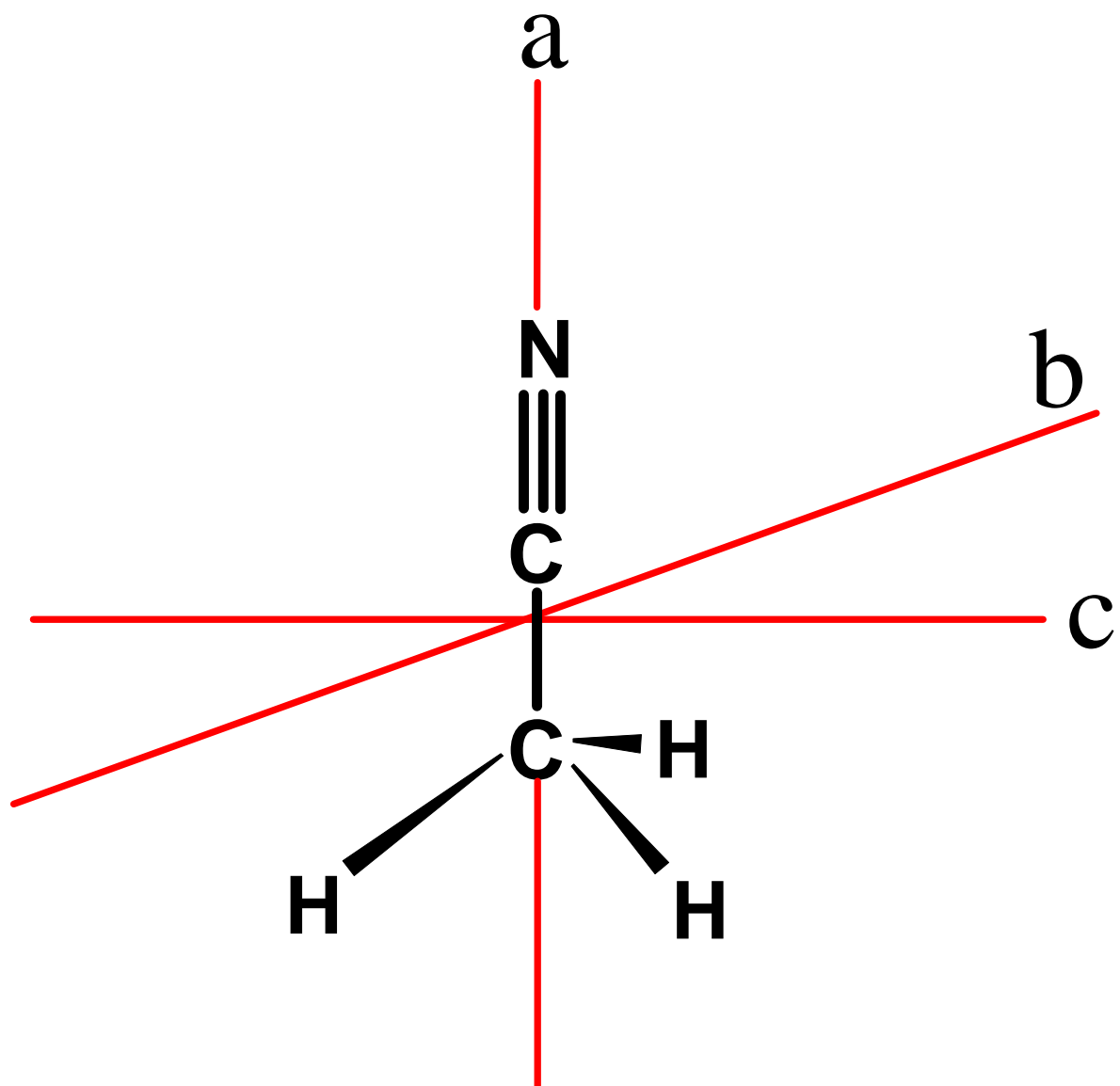


Figure 2

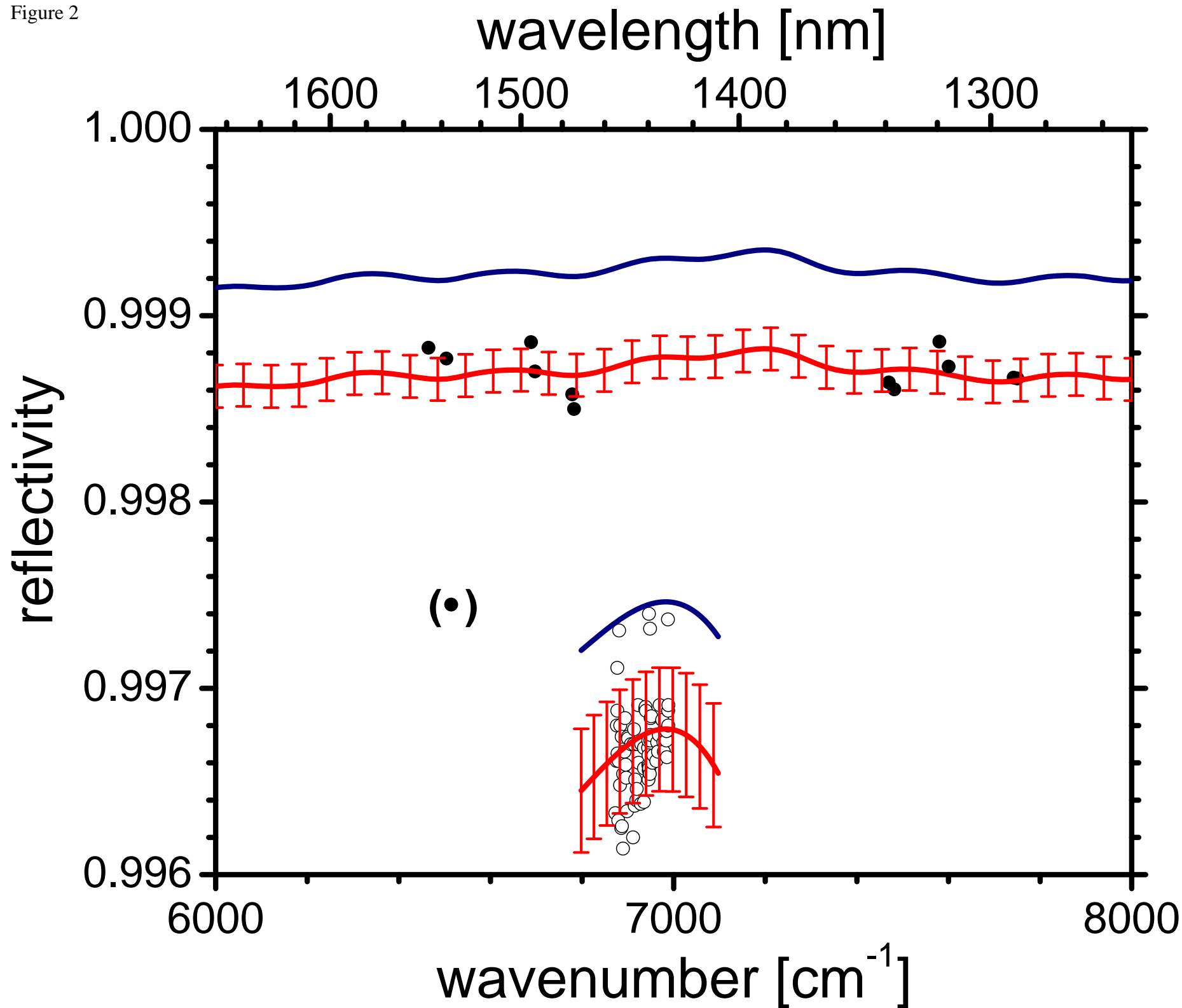


Figure 3

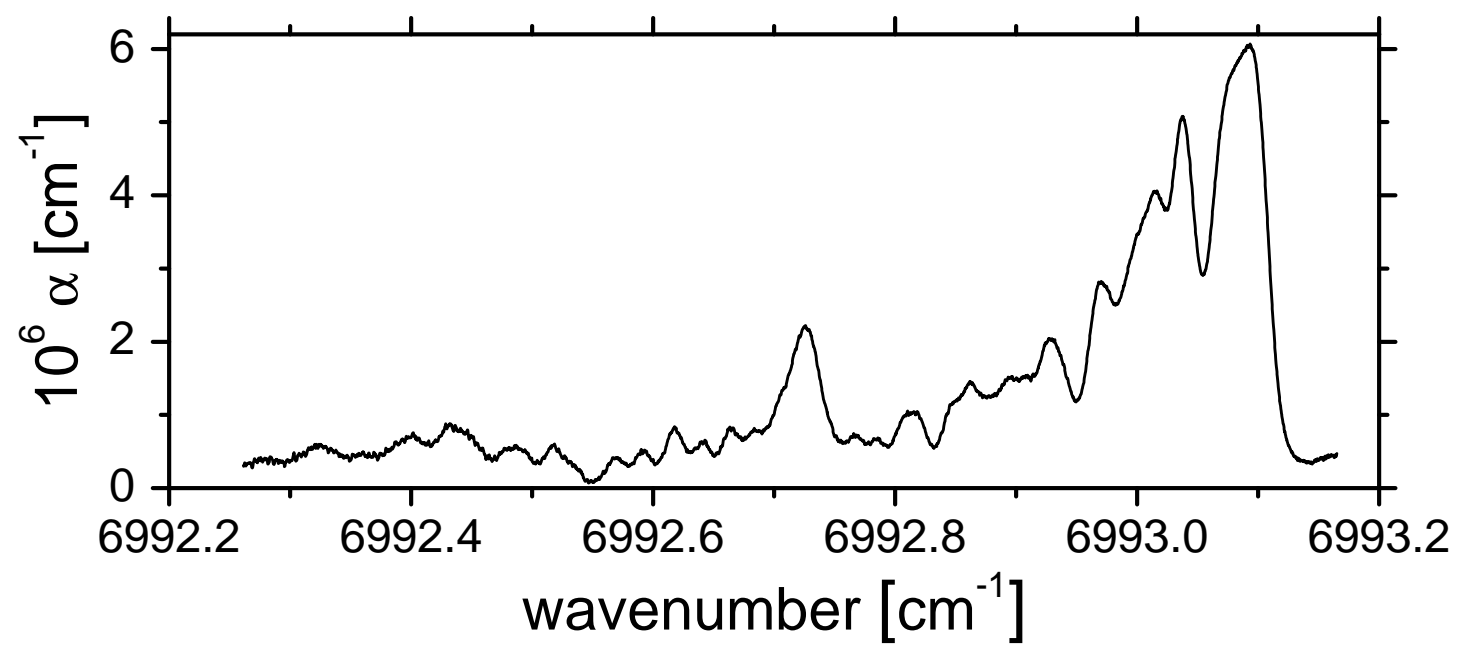


Figure 4

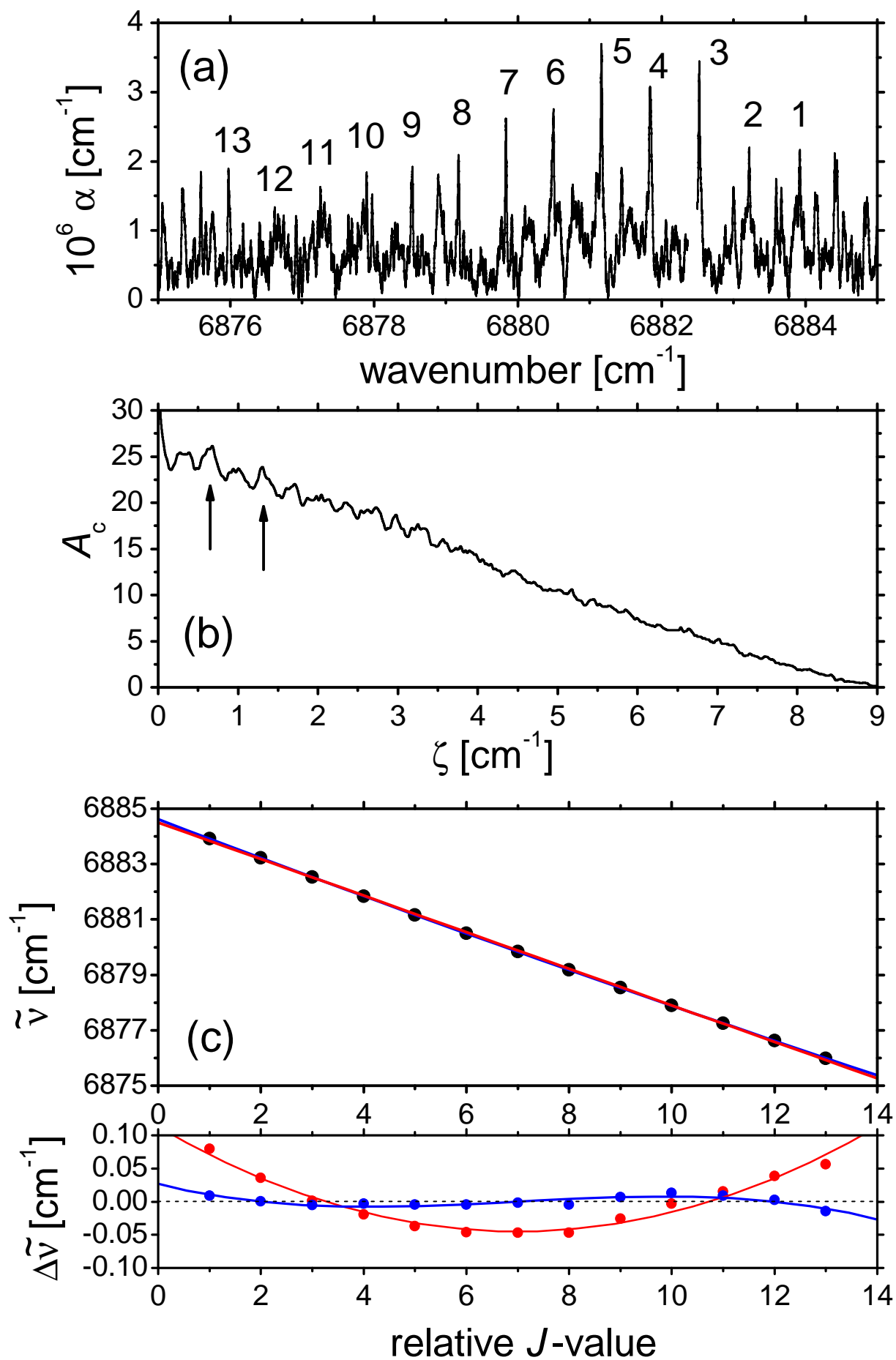


Figure 5

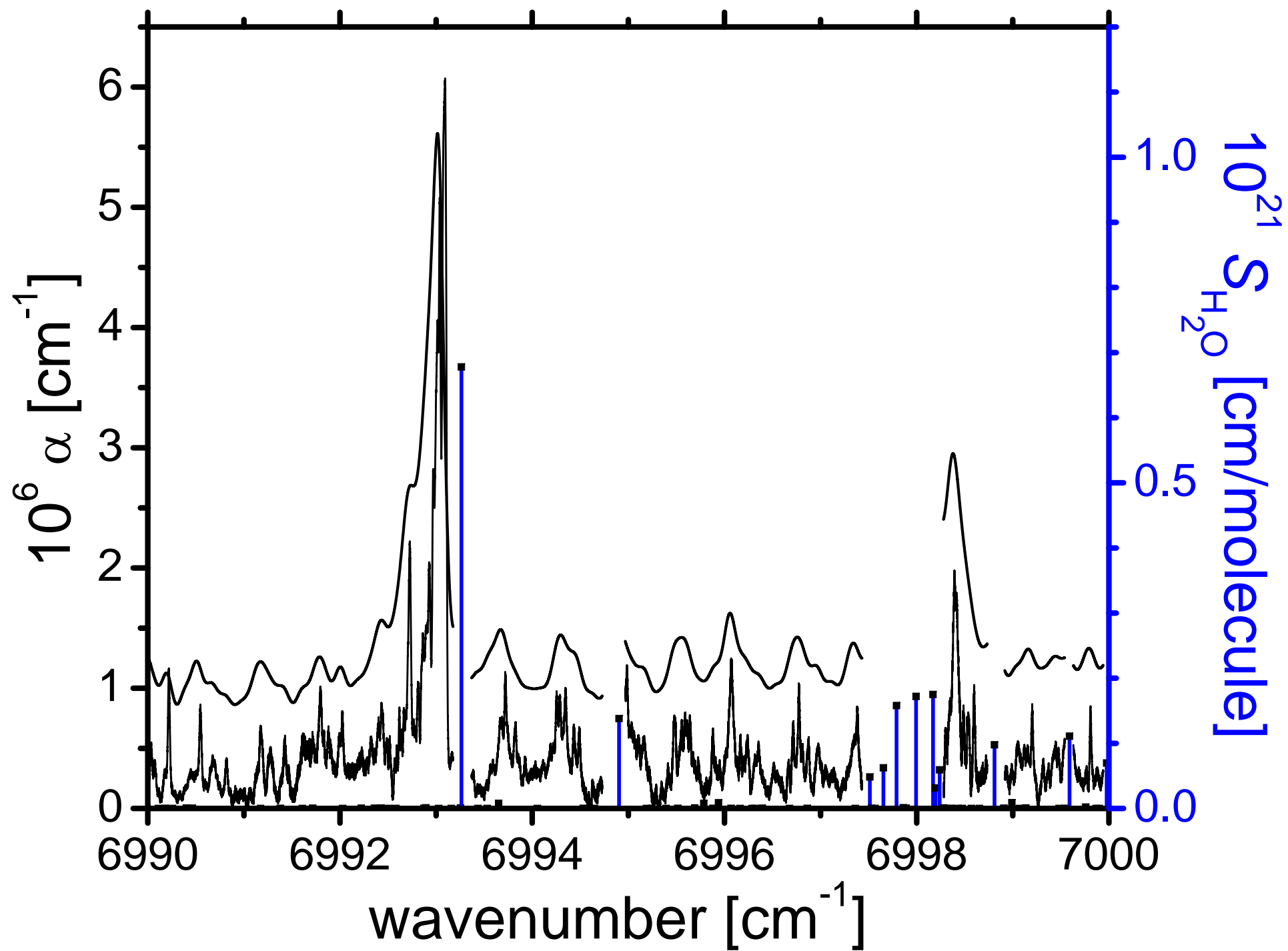




Figure 6

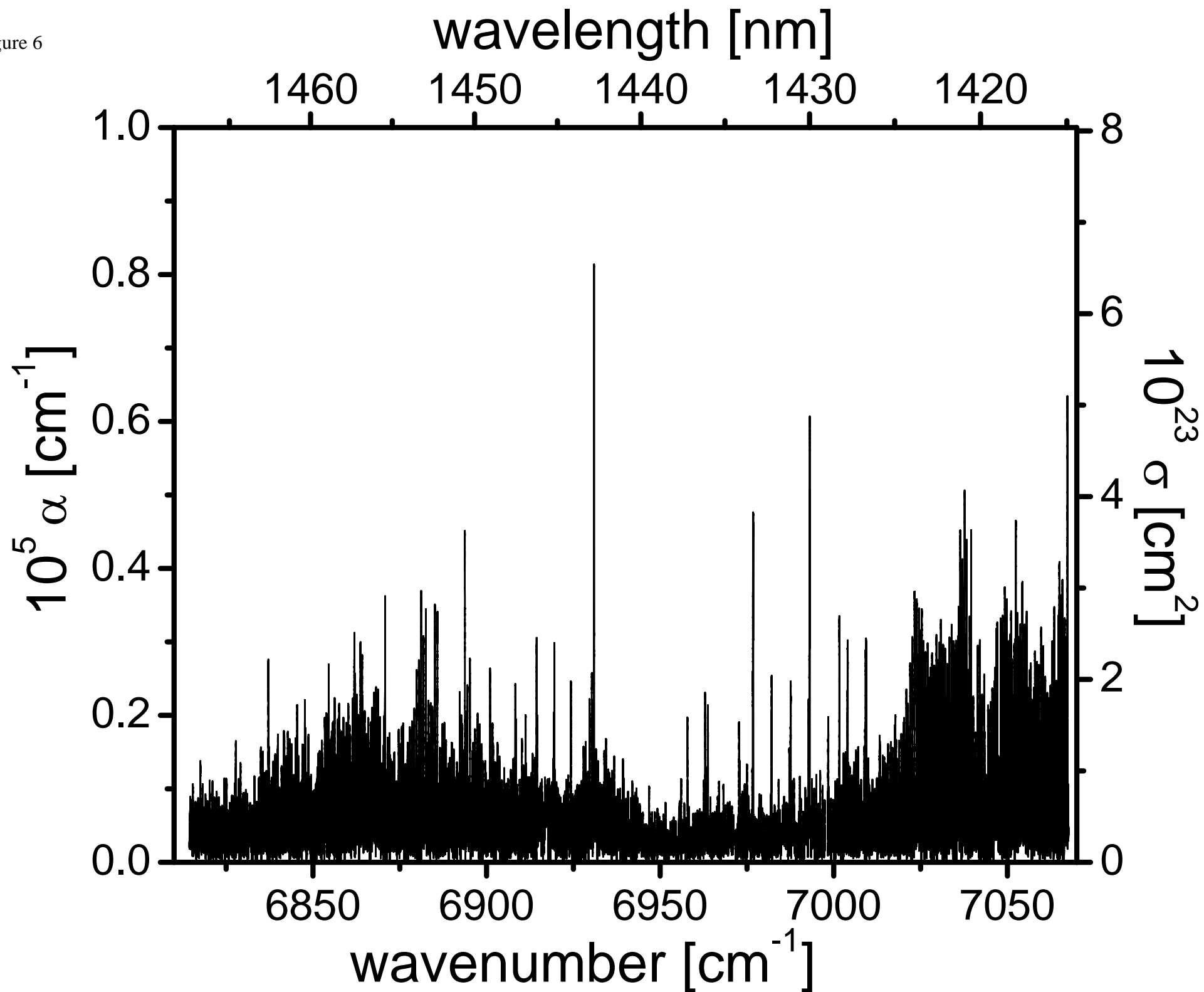


Figure 7

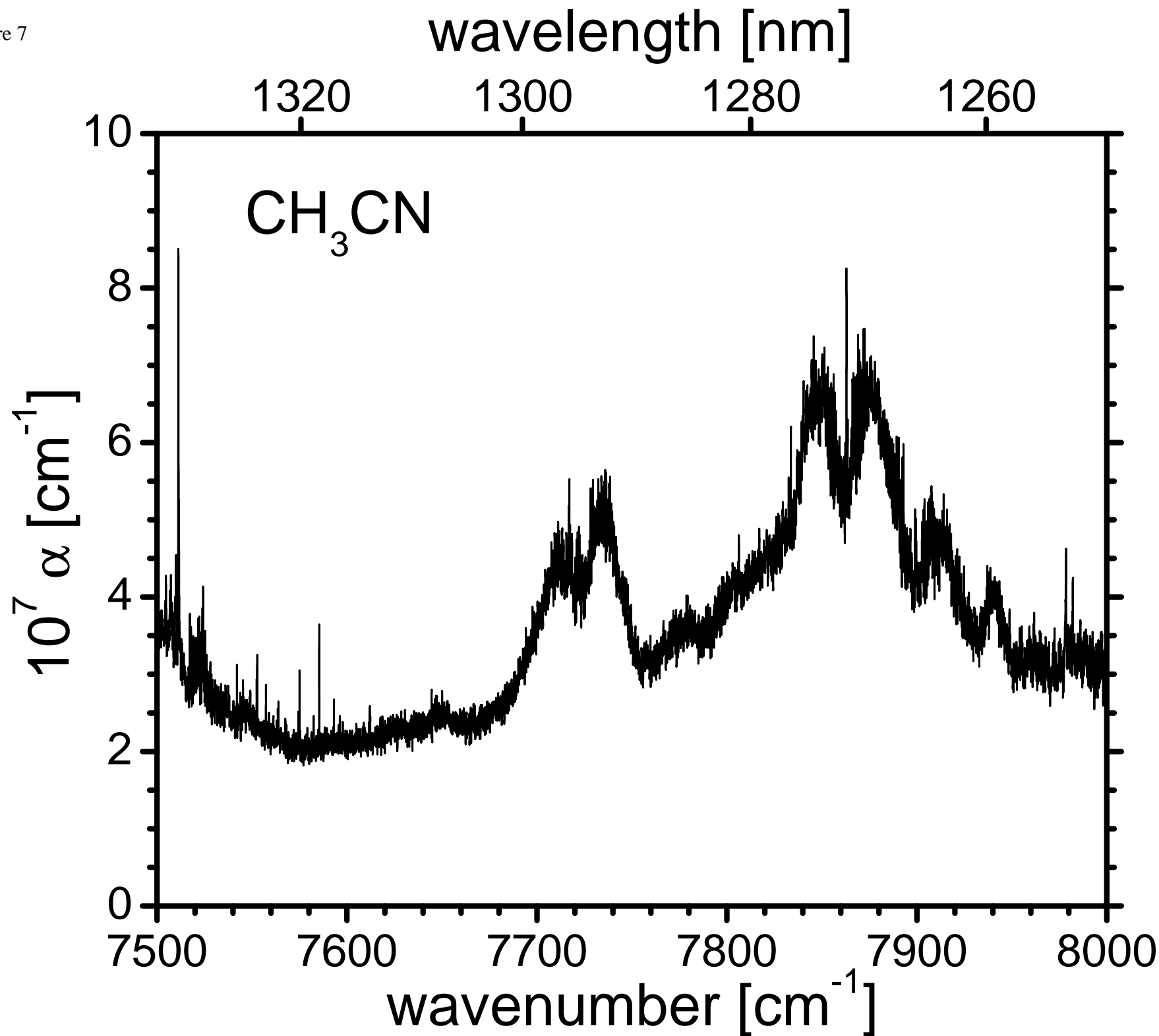


Figure 8

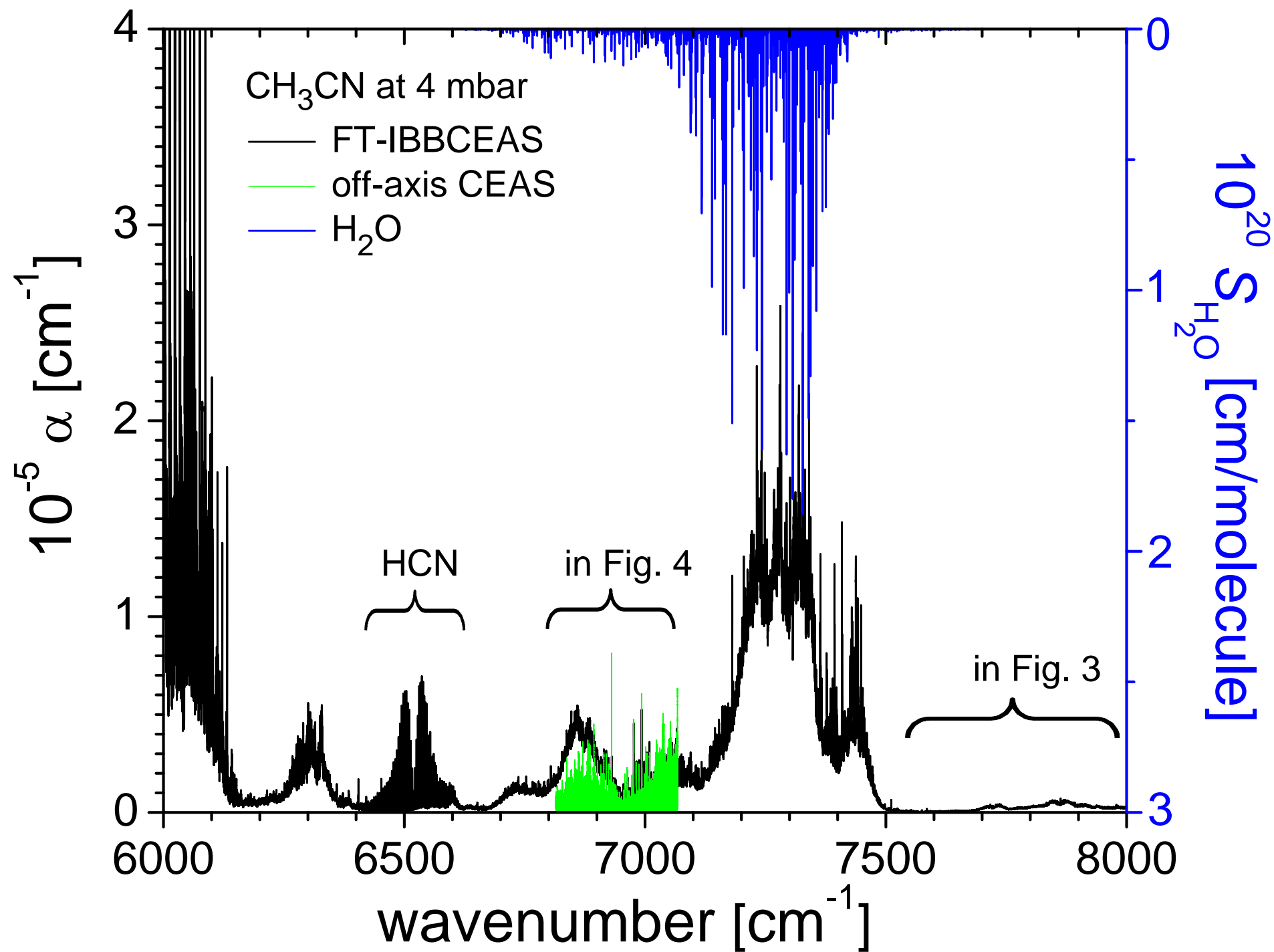


Figure 9

